### 'You Better Run'

Connecting low-energy Dark Matter searches with high-energy physics

Bradley J. Kavanagh LPTHE & IPhT (CEA/Saclay)

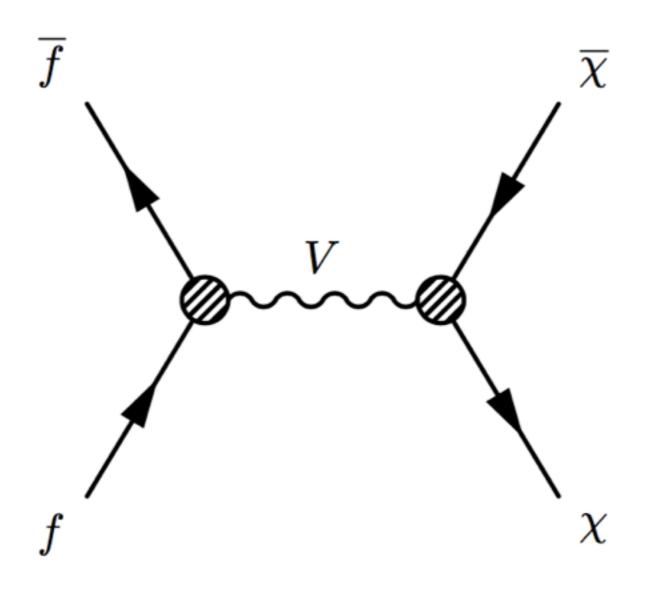


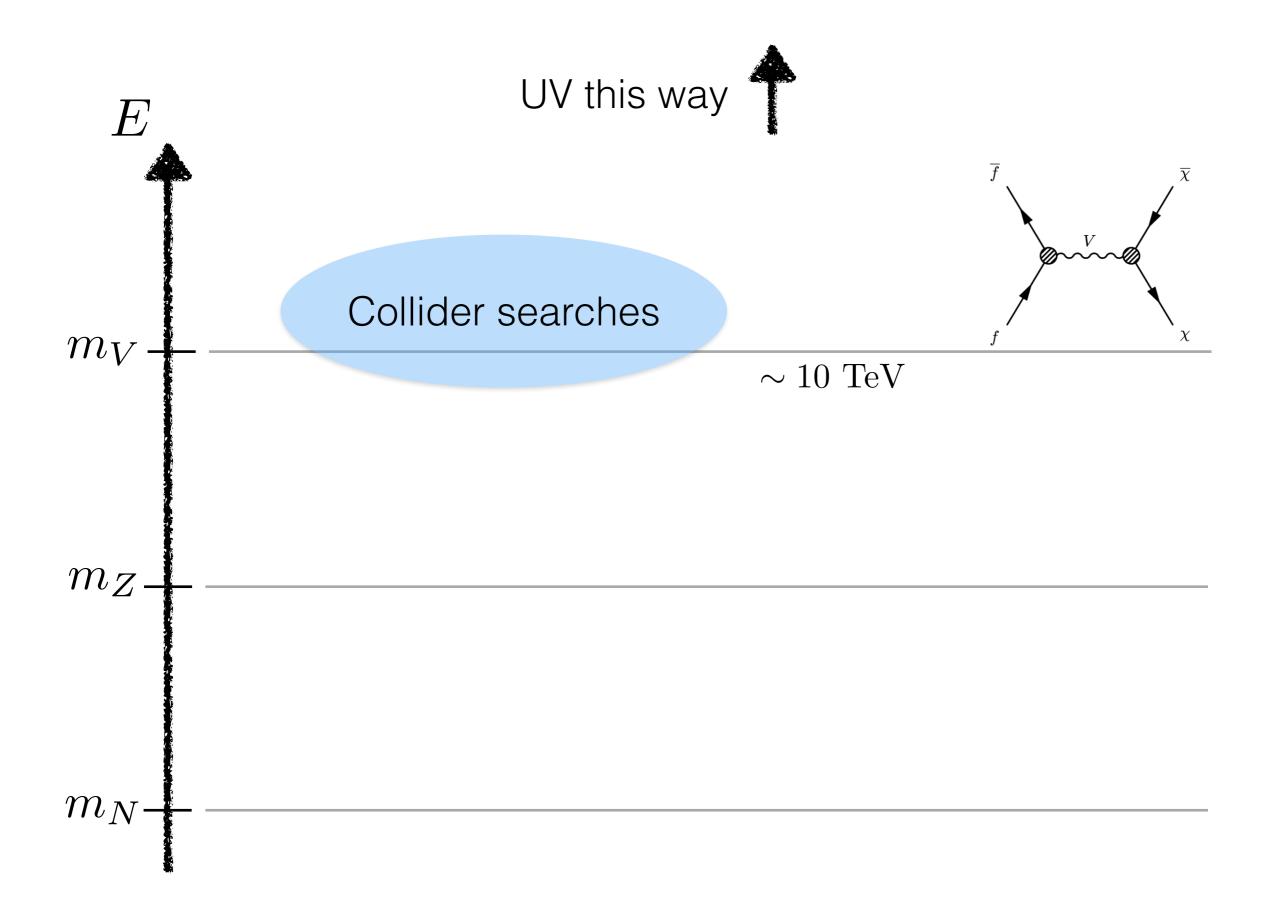


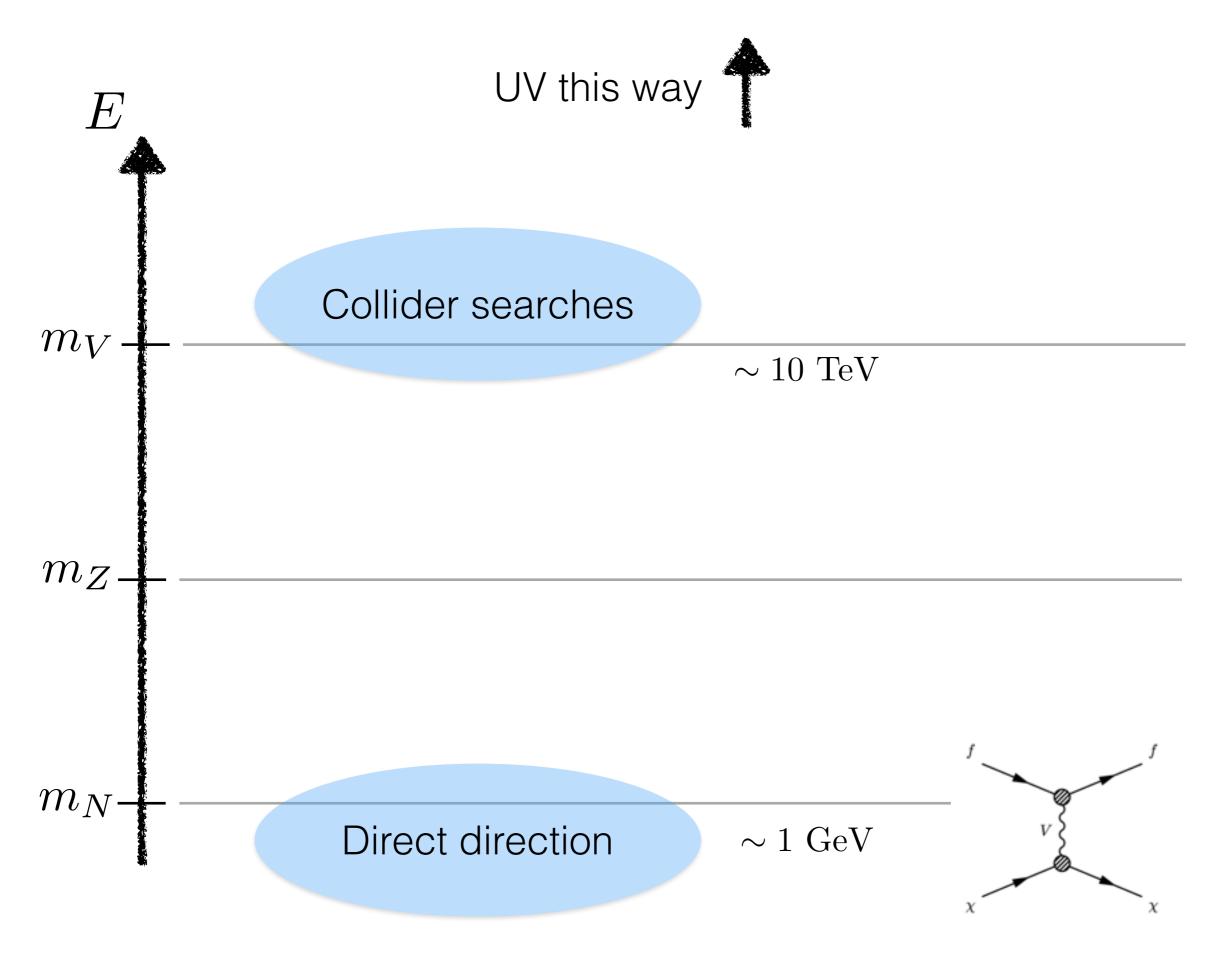


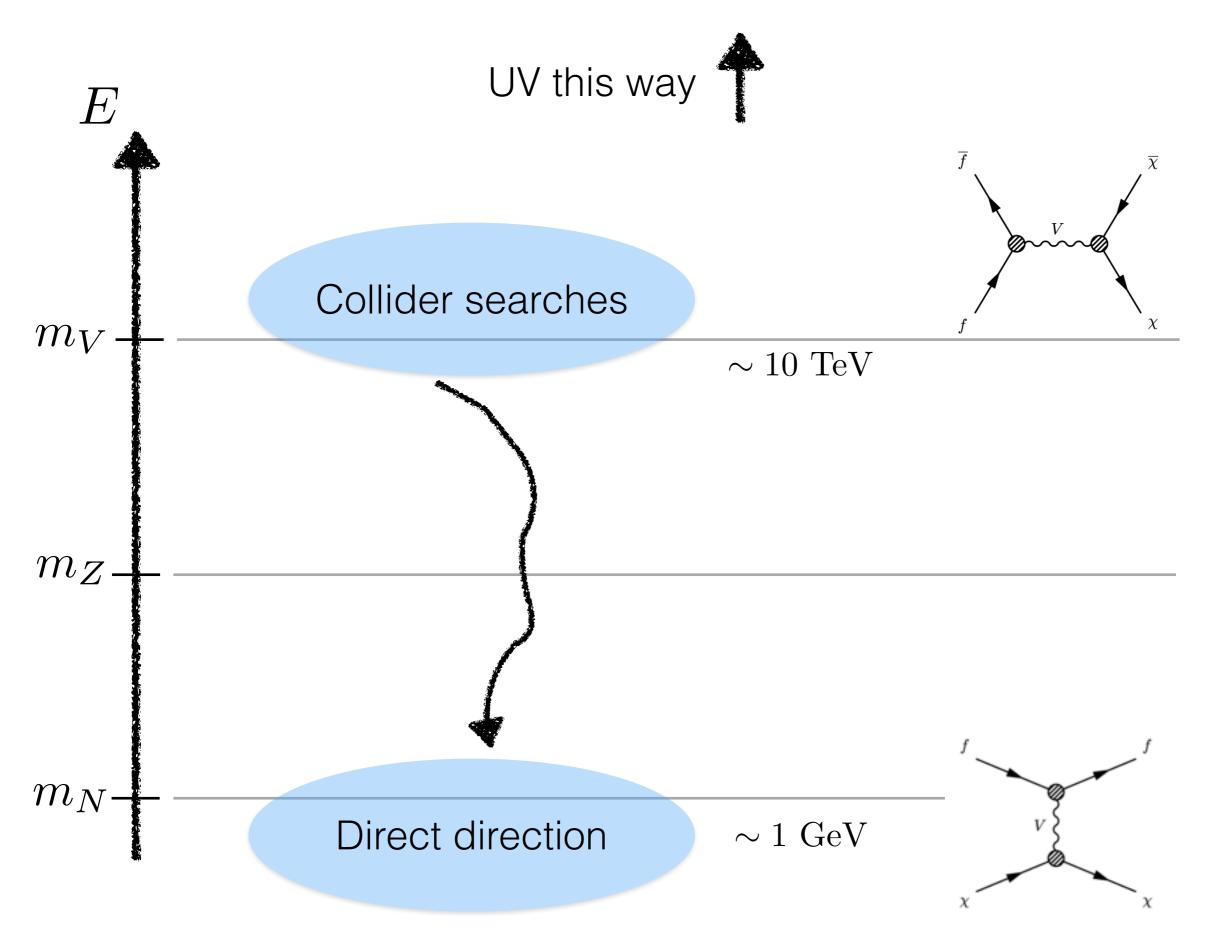


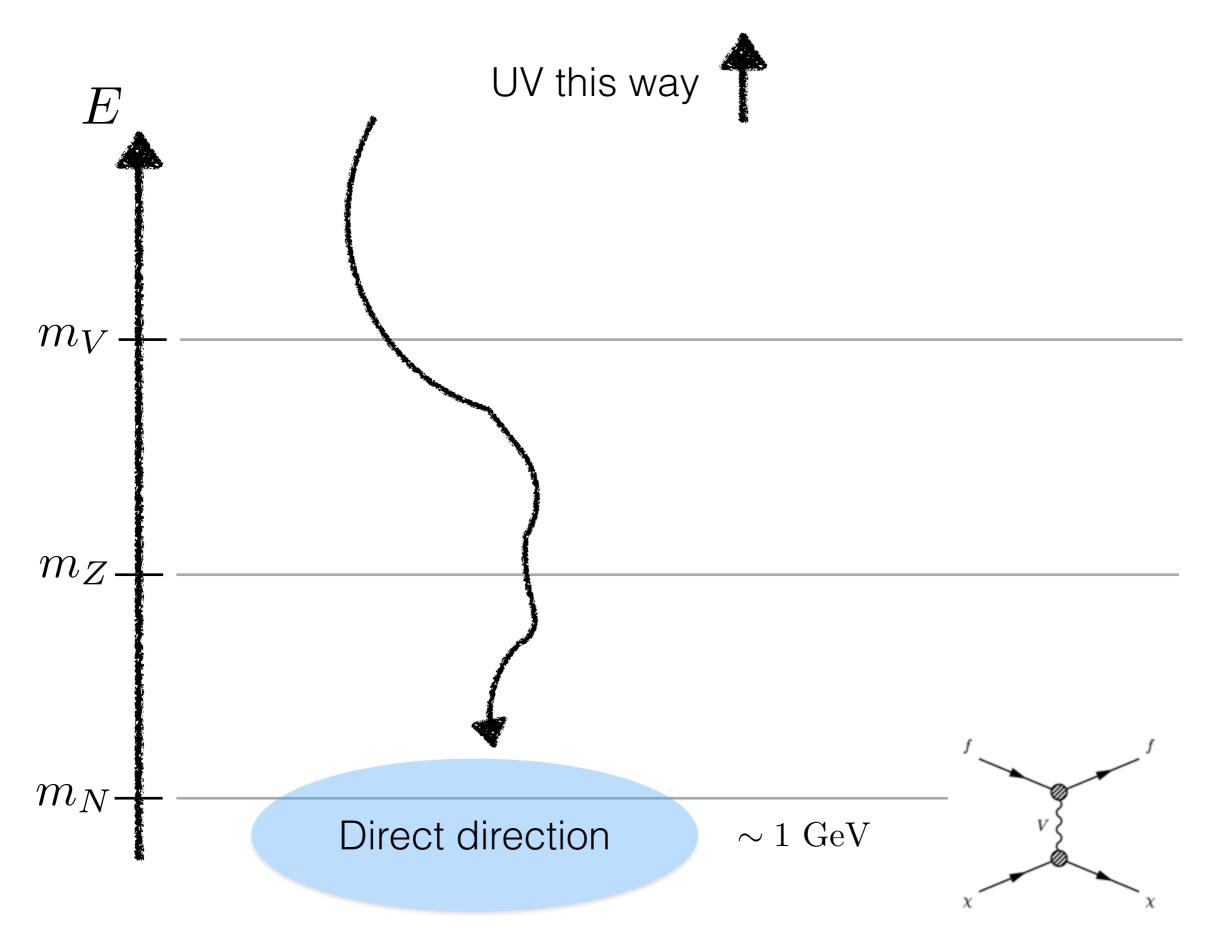




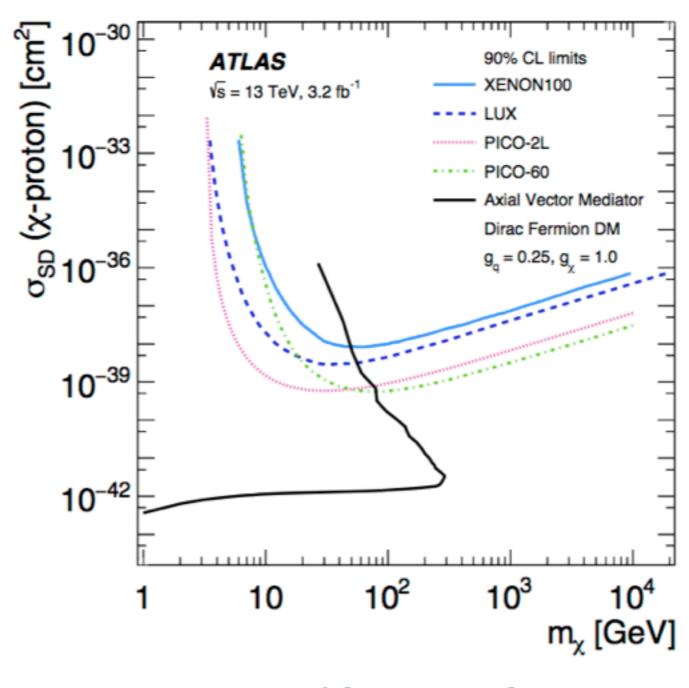








## Comparing different searches





ATLAS [1604.07773]

Zemeckis, Hanks (1994)

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#### **Outline**

### Simplified Models

De Simone, Jacques [1603.08002]

# RG effects in Simplified Models

Crivellin, D'Eramo, Procura [1402.1173]; D'Eramo, Procura [1411.3342]

## Direct detection constraints on Simplified Models

D'Eramo, Procura [1411.3342]; D'Eramo, BJK, Panci [1605.04917]

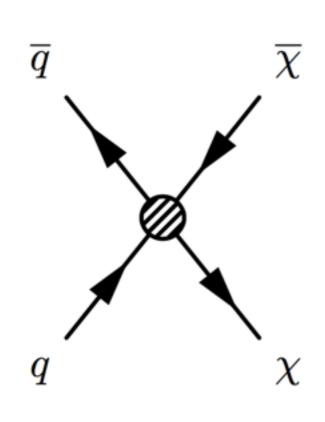
### Comparing DD and LHC searches

D'Eramo, BJK, Panci [1605.04917]

# Simplified Models

### **Effective Field Theory**

Assume mass of mediator is much larger than momentum transfer integrate out mediator to obtain a contact interaction

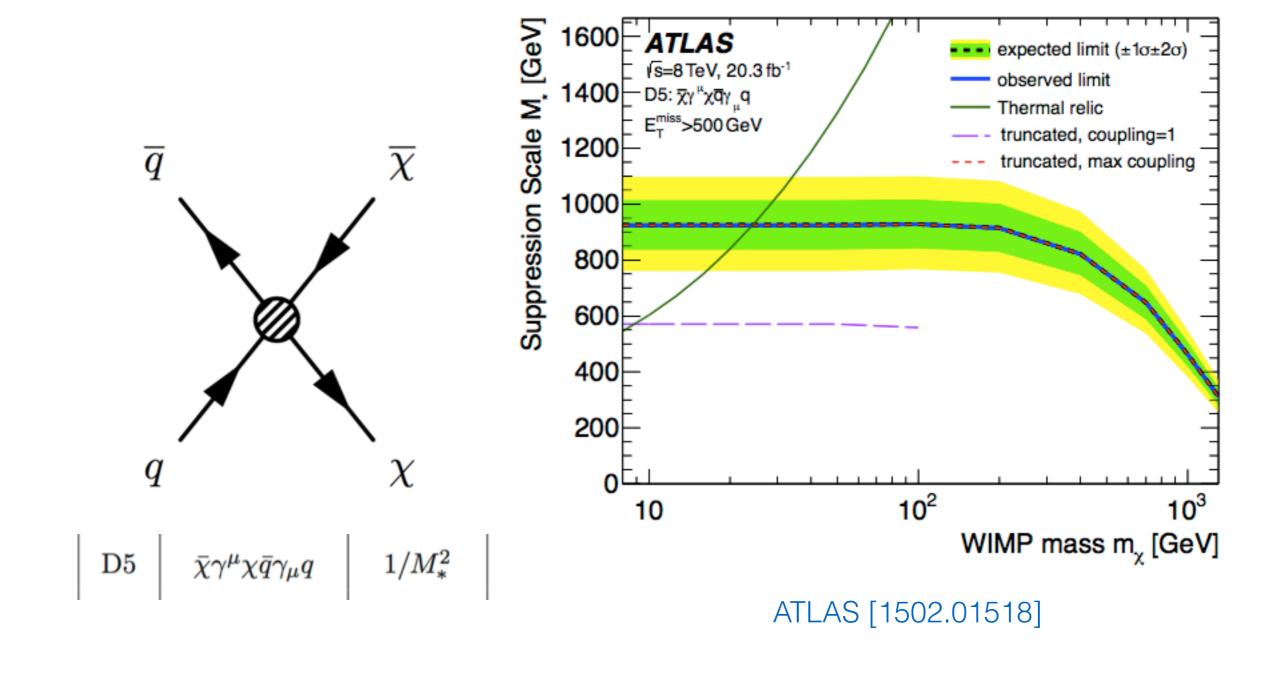


Name	Operator	Coefficient
D1	$ar{\chi}\chiar{q}q$	$m_q/M_*^3$
D2	$ar{\chi}\gamma^5\chiar{q}q$	$im_q/M_*^3$
D3	$ar{\chi}\chiar{q}\gamma^5q$	$im_q/M_*^3$
D4	$\bar{\chi}\gamma^5\chi \bar{q}\gamma^5q$	$m_q/M_*^3$
D5	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D6	$ar{\chi}\gamma^{\mu}\gamma^5\chiar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D7	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$
D8	$ar{\chi}\gamma^{\mu}\gamma^5\chiar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$
D9	$\bar{\chi}\sigma^{\mu u}\chiar{q}\sigma_{\mu u}q$	$1/M_{*}^{2}$
D10	$ar{\chi}\sigma_{\mu u}\gamma^5\chiar{q}\sigma_{lphaeta}q$	$i/M_*^2$

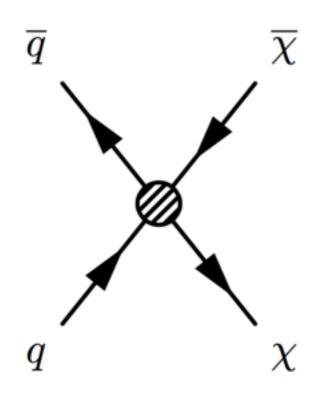
Goodman et al. [1008.1783]

Only have to deal with two parameters:  $m_\chi, \Lambda$ 

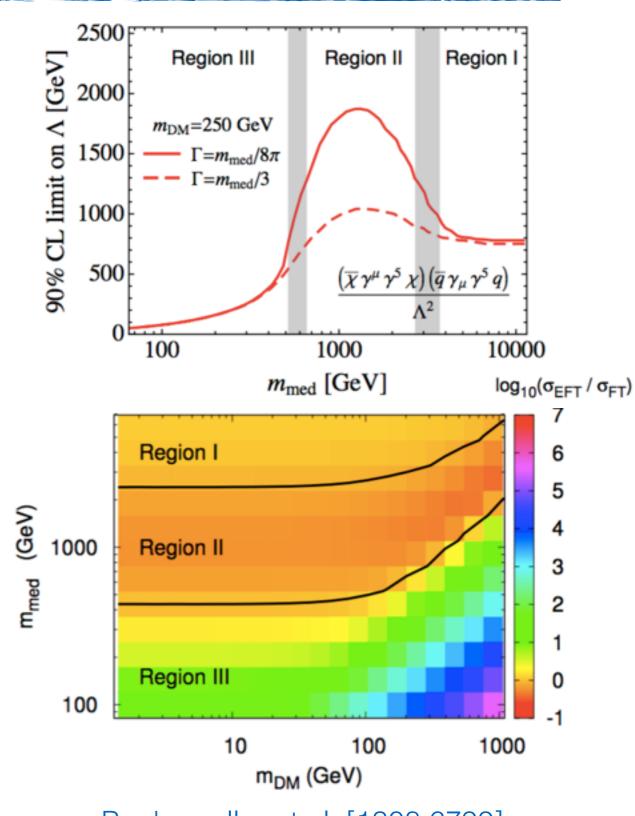
### **Limits on EFT**



### The problem with EFTs



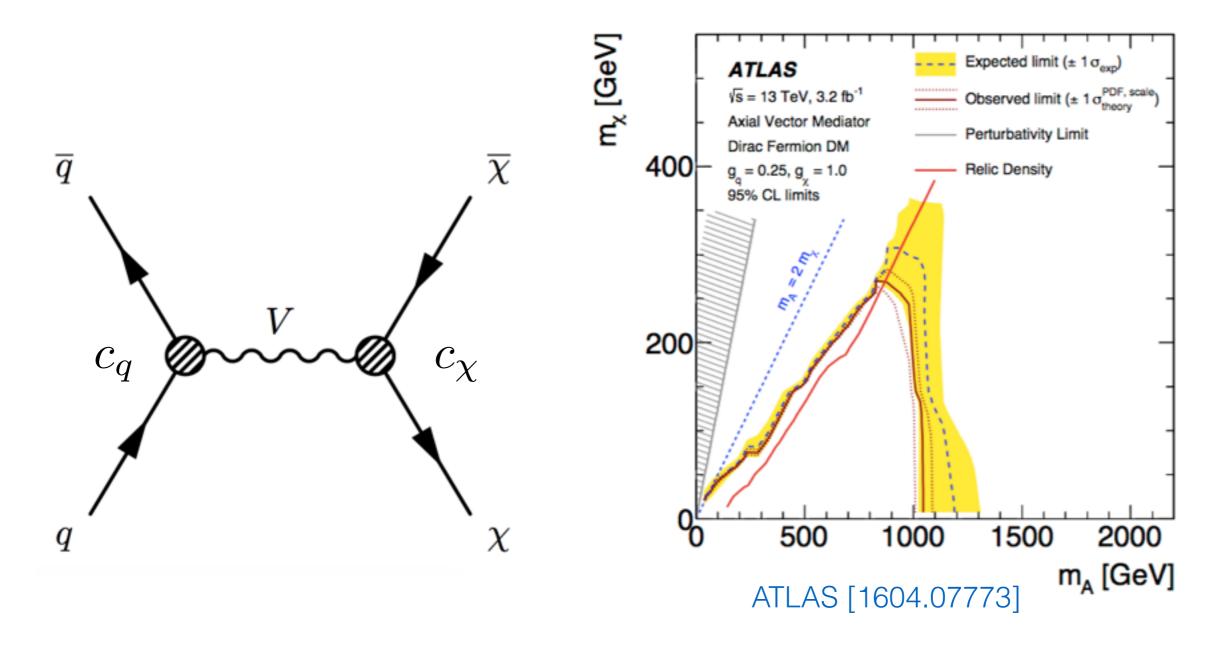
Frandsen et al. [1204.3839], Buchmueller et al. [1407.8257], Malik et al. [1409.4075], Abdallah et al. [1506.03116], and many others...



Buchmueller et al. [1308.6799]

### Simplified Models to the rescue

Review: De Simone, Jacques [1603.08002]



Now have to deal with more parameters:  $c_q, c_\chi, m_\chi, m_V$ 

# Our Simplified Model

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} + \mathcal{L}_{V} + J_{DM}^{\mu} V_{\mu} + J_{SM}^{\mu} V_{\mu}$$

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} + \mathcal{L}_{V} + J_{DM}^{\mu} V_{\mu} + J_{SM}^{\mu} V_{\mu}$$

$$\mathcal{L}_{\mathrm{DM}} \sim \begin{cases} \left| \partial_{\mu} \phi \right|^{2} - m_{\phi}^{2} \left| \phi \right|^{2} & \text{complex scalar DM} \\ \overline{\chi} \left( i \partial \hspace{-0.1cm} / - m_{\chi} \right) \chi & \text{fermion DM} \end{cases}$$

$$\mathcal{L} = \mathcal{L}_{\mathrm{SM}} + \mathcal{L}_{\mathrm{DM}} + \mathcal{L}_{\mathrm{V}} + J_{\mathrm{DM}}^{\mu} V_{\mu} + J_{\mathrm{SM}}^{\mu} V_{\mu}$$

$$\mathcal{L}_{V} = -\frac{1}{4} V^{\mu\nu} V_{\mu\nu} + \frac{1}{2} m_V^2 V^{\mu} V_{\mu}$$

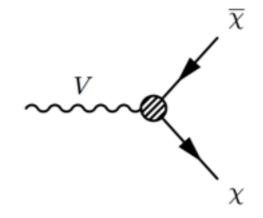
Massive spin-1 mediator V

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{DM} + \mathcal{L}_{V} + J_{DM}^{\mu} V_{\mu} + J_{SM}^{\mu} V_{\mu}$$



$$J_{\rm DM}^{\mu} \sim \begin{cases} c_{\phi} \phi^{\dagger} \overleftrightarrow{\partial}_{\mu} \phi \\ c_{\chi V} \overline{\chi} \gamma^{\mu} \chi + c_{\chi A} \overline{\chi} \gamma^{\mu} \gamma^{5} \chi \end{cases}$$

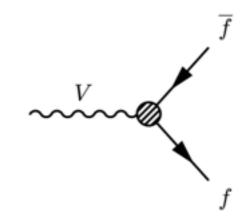
complex scalar DM fermion DM



$$\mathcal{L} = \mathcal{L}_{\mathrm{SM}} + \mathcal{L}_{\mathrm{DM}} + \mathcal{L}_{\mathrm{V}} + J_{\mathrm{DM}}^{\mu} V_{\mu} + J_{\mathrm{SM}}^{\mu} V_{\mu}$$

$$J_{\rm SM}^{\mu} = \sum_{i=1}^{3} \left[ c_q^{(i)} \ \overline{q_L^i} \gamma^{\mu} q_L^i + c_u^{(i)} \ \overline{u_R^i} \gamma^{\mu} u_R^i + c_d^{(i)} \ \overline{d_R^i} \gamma^{\mu} d_R^i + c_l^{(i)} \ \overline{l_L^i} \gamma^{\mu} l_L^i + c_e^{(i)} \ \overline{e_R^i} \gamma^{\mu} e_R^i \right]$$

15 independent, SU(2)<sub>L</sub> x U(1)<sub>Y</sub> gauge-invariant couplings

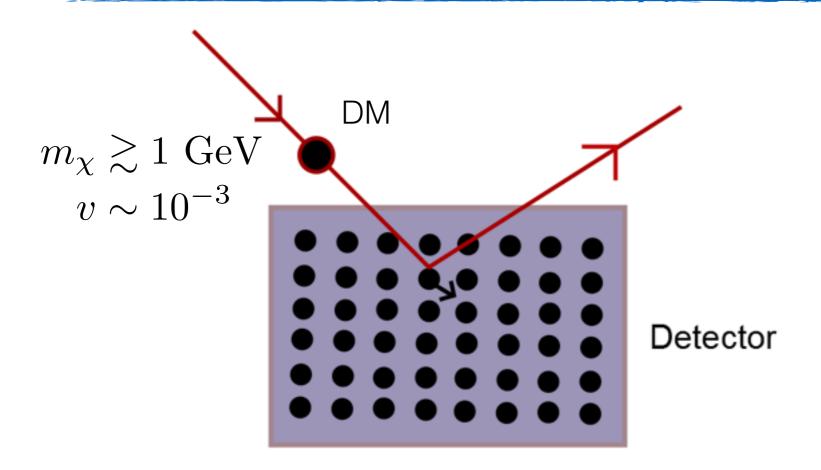


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Calculating the direct detection rate

### **Direct detection**



Look for low energy - O(keV) - recoils of detector nuclei

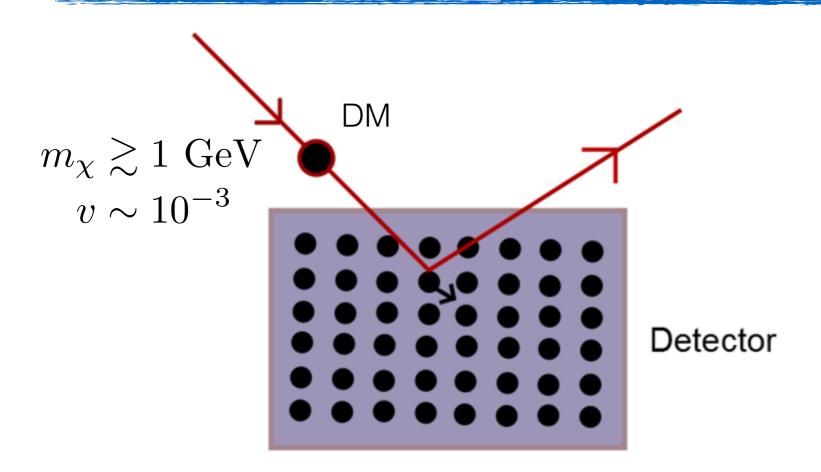
Rate driven by coupling of DM to light quarks (u, d, s):

$$\overline{\chi}\gamma^{\mu}\chi \,\overline{q}\gamma_{\mu}q \qquad \overline{\chi}\gamma^{\mu}\chi \,\overline{q}\gamma_{\mu}\gamma^{5}q 
\overline{\chi}\gamma^{\mu}\gamma^{5}\chi \,\overline{q}\gamma_{\mu}q \qquad \overline{\chi}\gamma^{\mu}\gamma^{5}\chi \,\overline{q}\gamma_{\mu}\gamma^{5}q$$

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### **Direct detection**



Look for low energy - O(keV) - recoils of detector nuclei

Rate driven by coupling of DM to light quarks (u, d, s):

Standard SI

$$\overline{\chi}\gamma^{\mu}\chi \ \overline{q}\gamma_{\mu}q$$

$$\overline{\chi}\gamma^{\mu}\chi \ \overline{q}\gamma_{\mu}\gamma^5 q$$

$$\overline{\chi}\gamma^{\mu}\gamma^5\chi \ \overline{q}\gamma_{\mu}q$$

$$\overline{\chi}\gamma^{\mu}\gamma^{5}\chi \ \overline{q}\gamma_{\mu}\gamma^{5}q$$

Standard SD

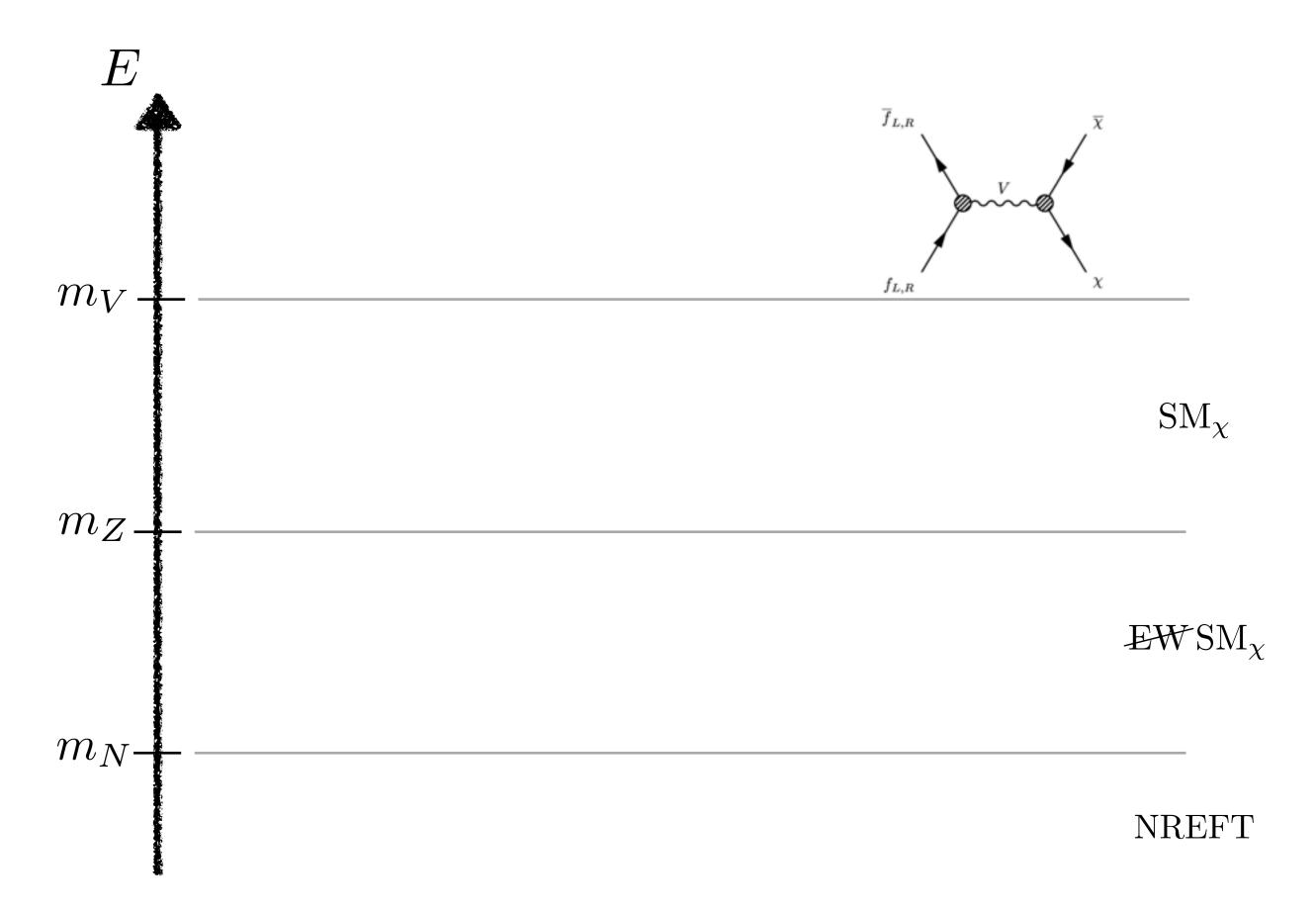
## Connecting high and low scales

Define couplings at high energy scale (mediator mass), but need to calculate direct detection rate at low energy

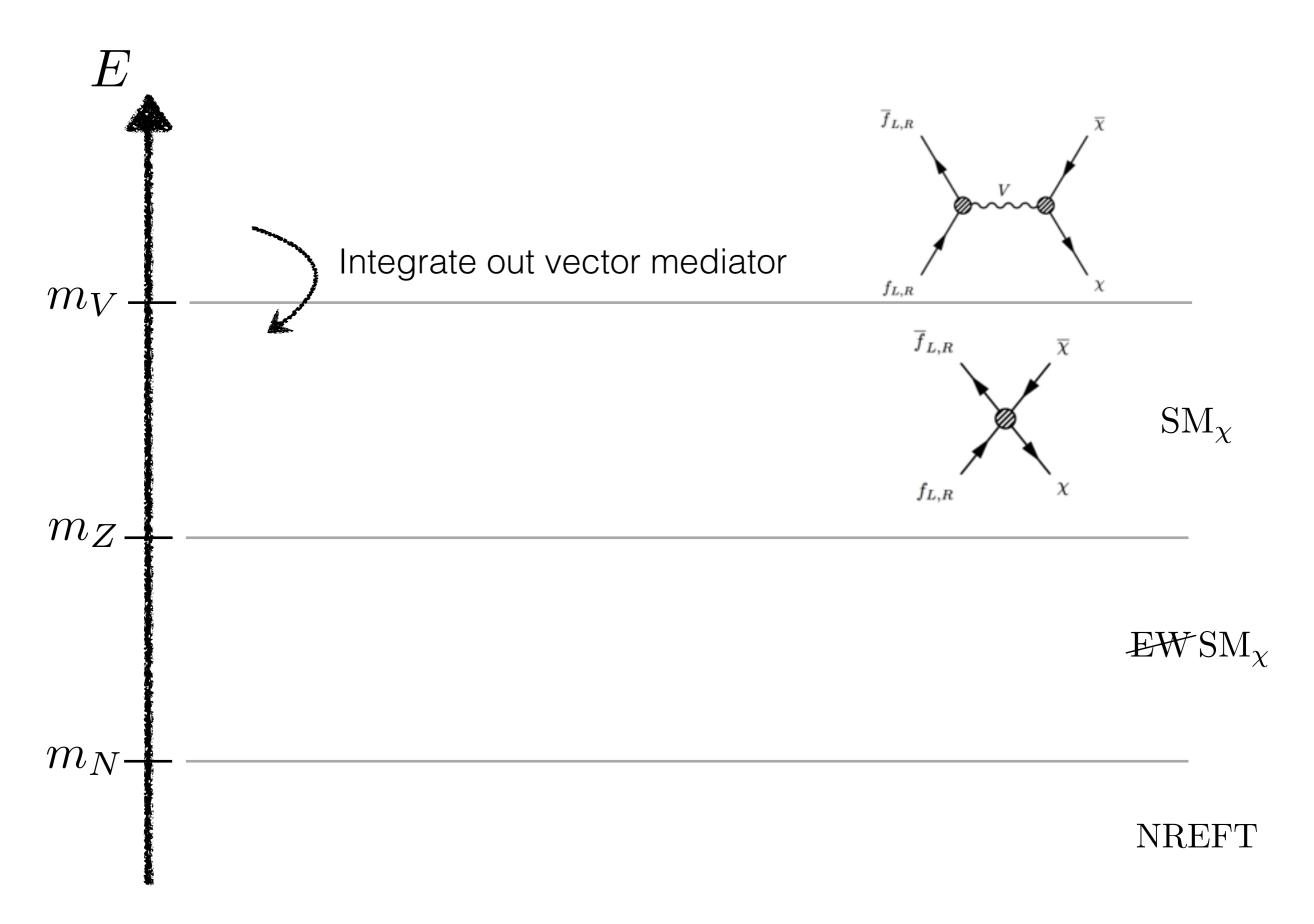
Running can change the DD rate by orders of magnitude. Examples in specific models:

Kopp et al. [0907.3159], Frandsen et al. [1207.3971], Haisch, Kahlhoefer [1302.4454], Kopp et al. [1401.6457], Crivellin, Haisch [1408.5046]

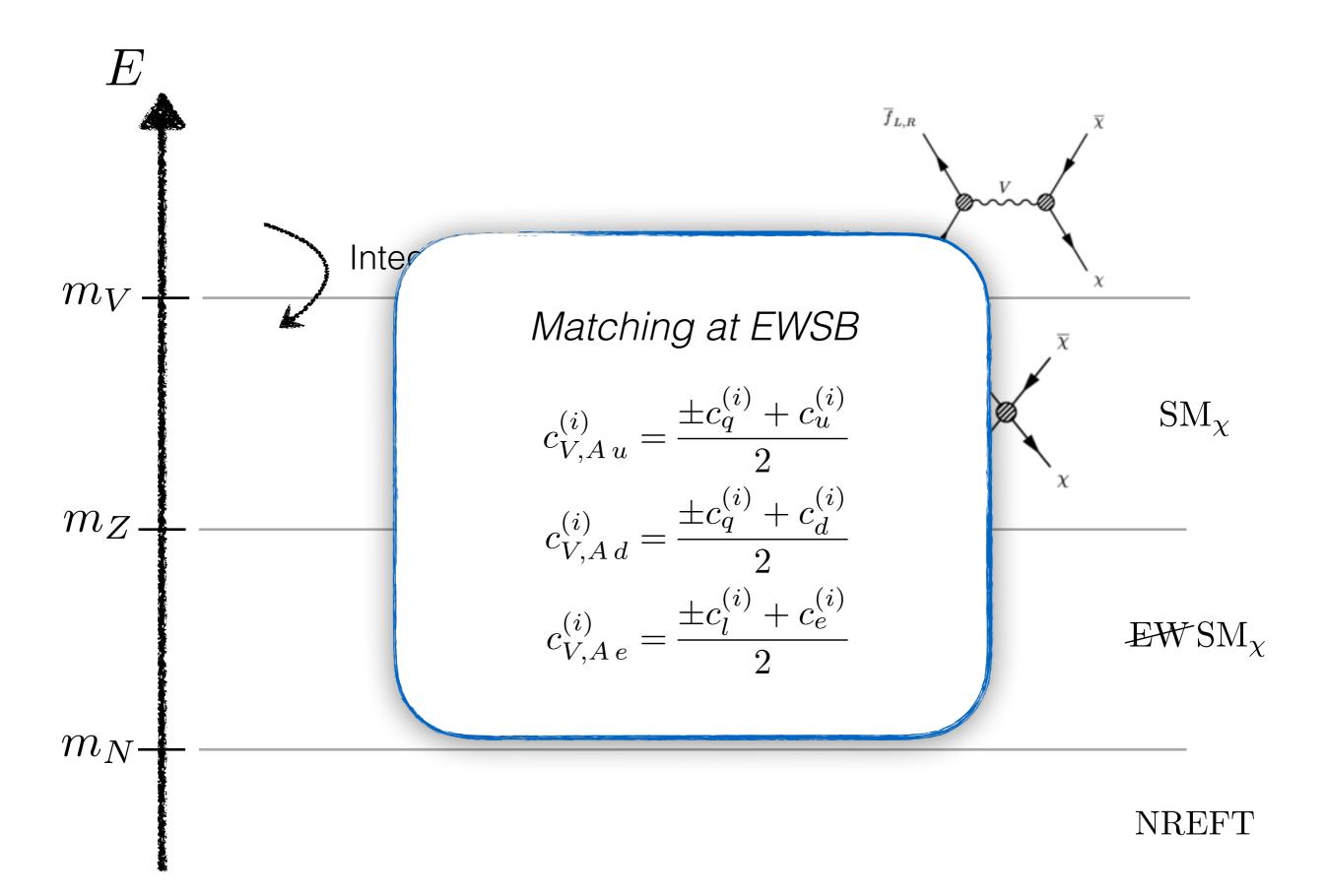
- Use EFT techniques and RG flow to study the effects for general interactions
- Include all relevant DD interactions (not just naive 'leading order')



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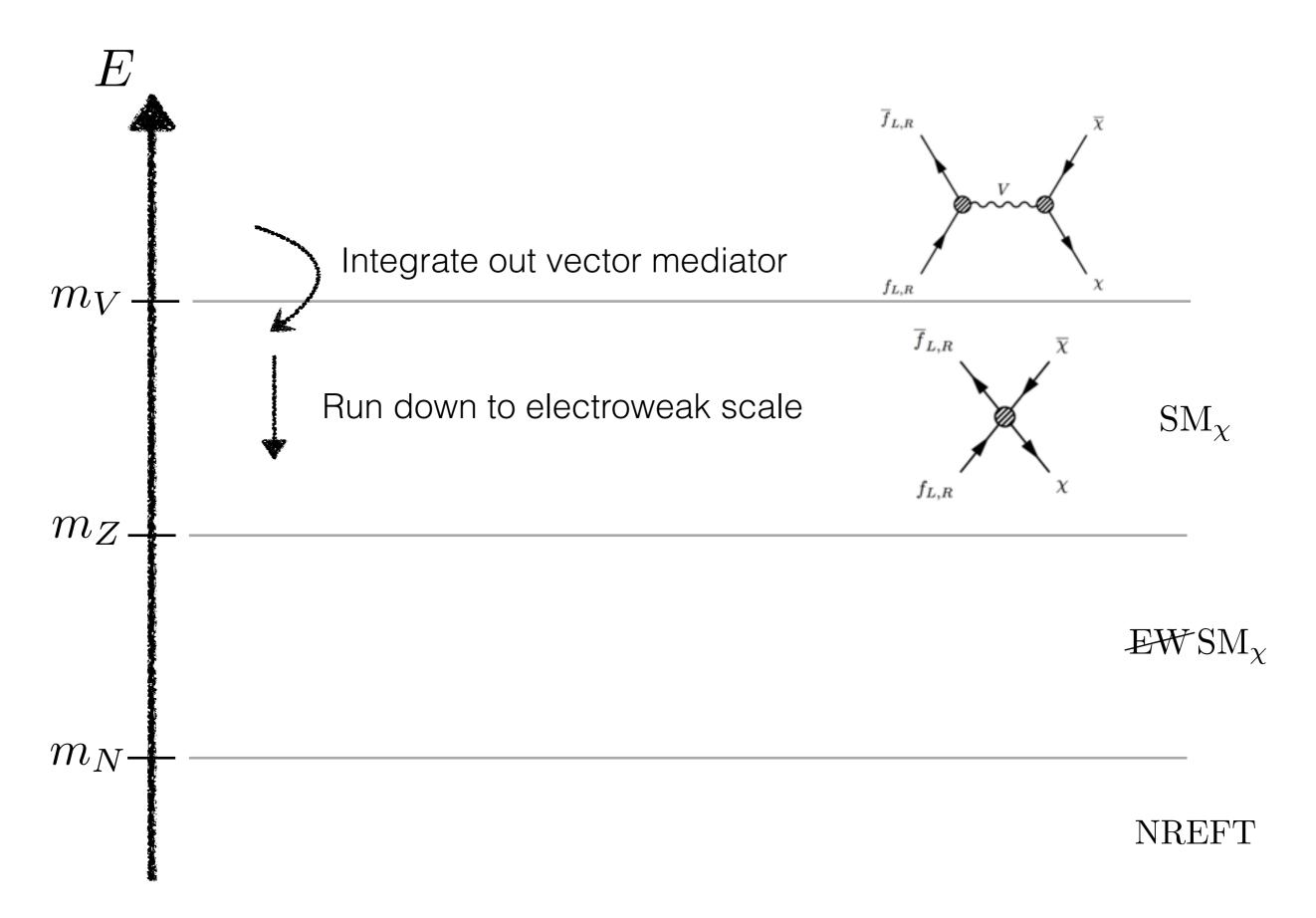


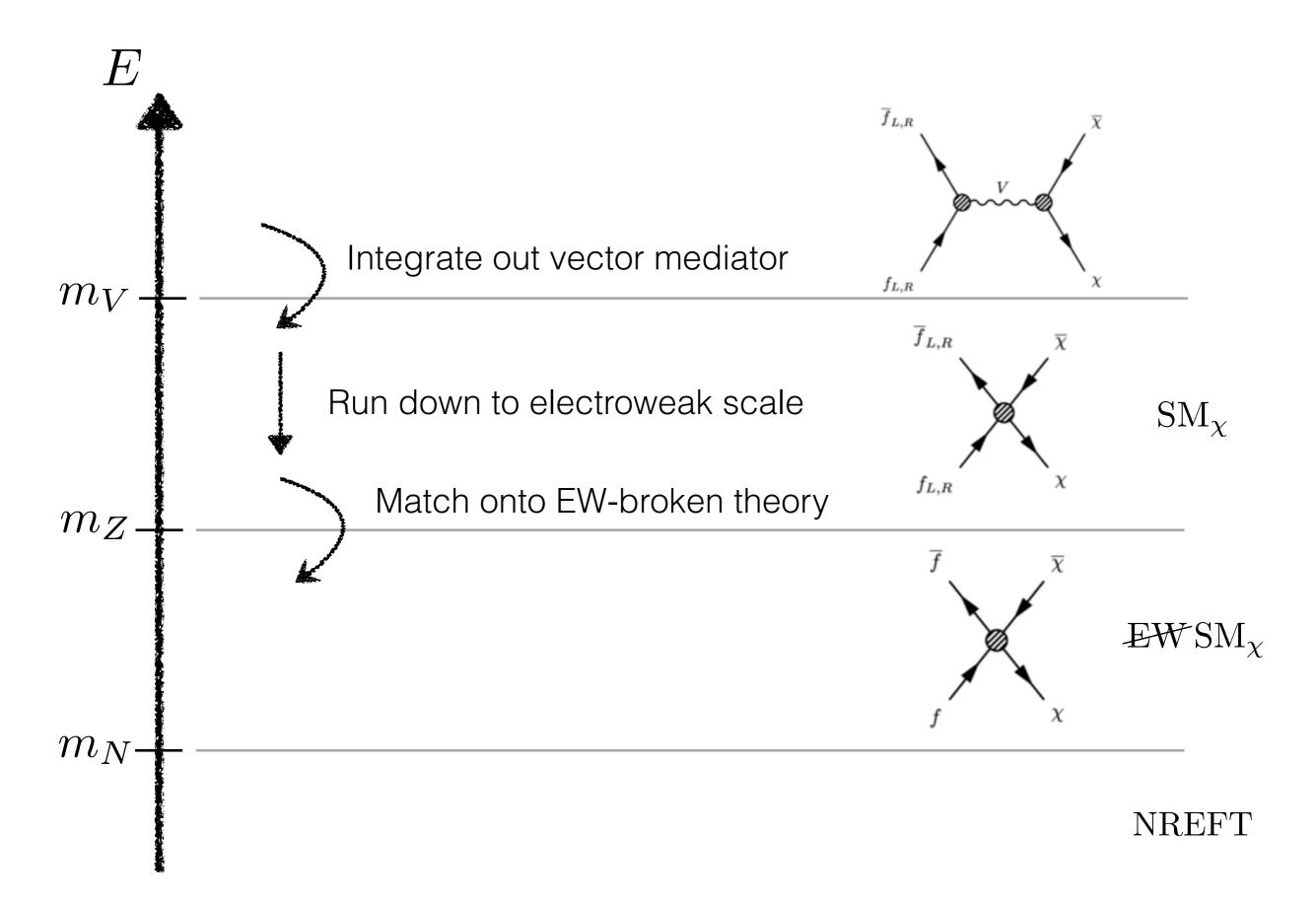
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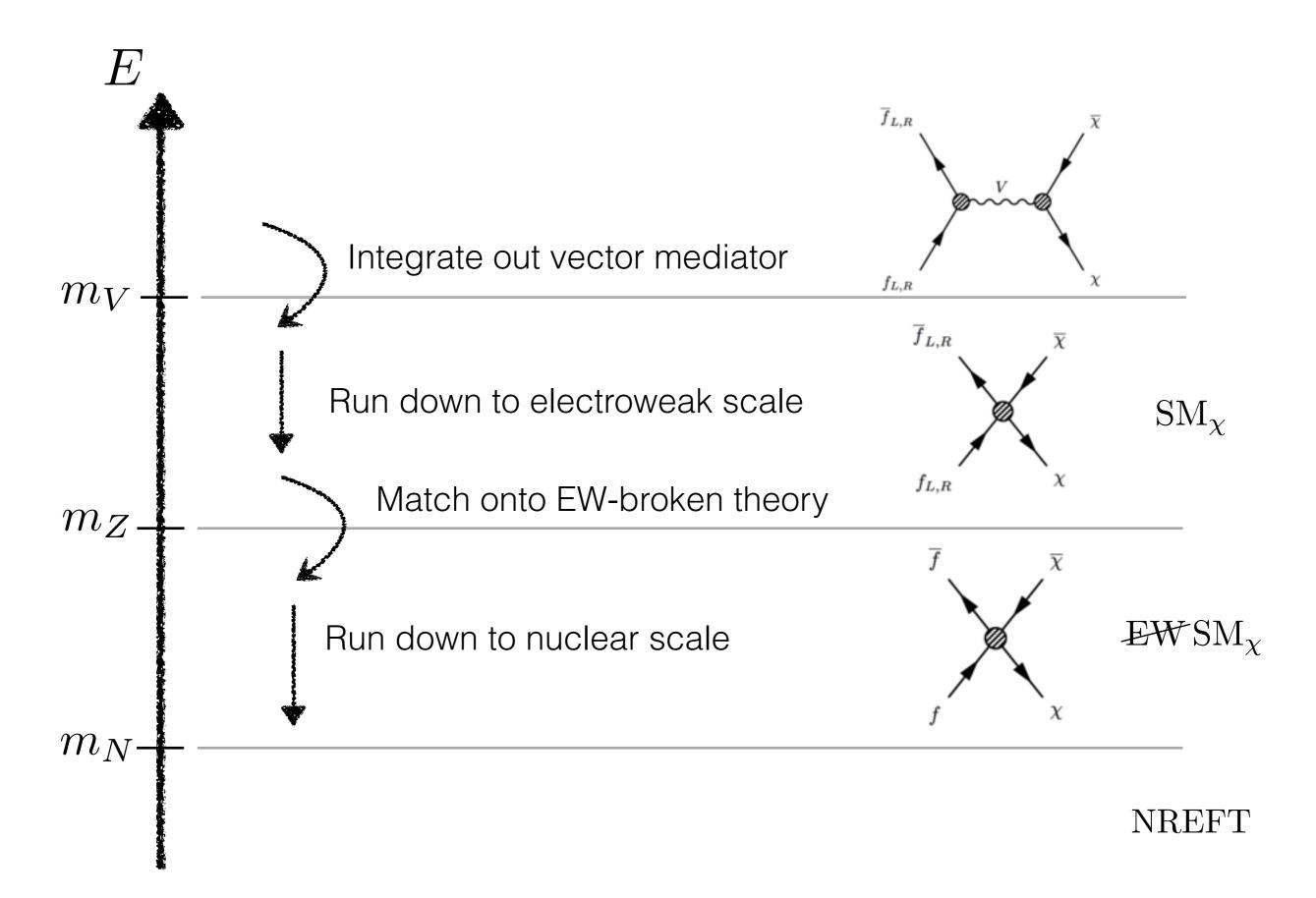


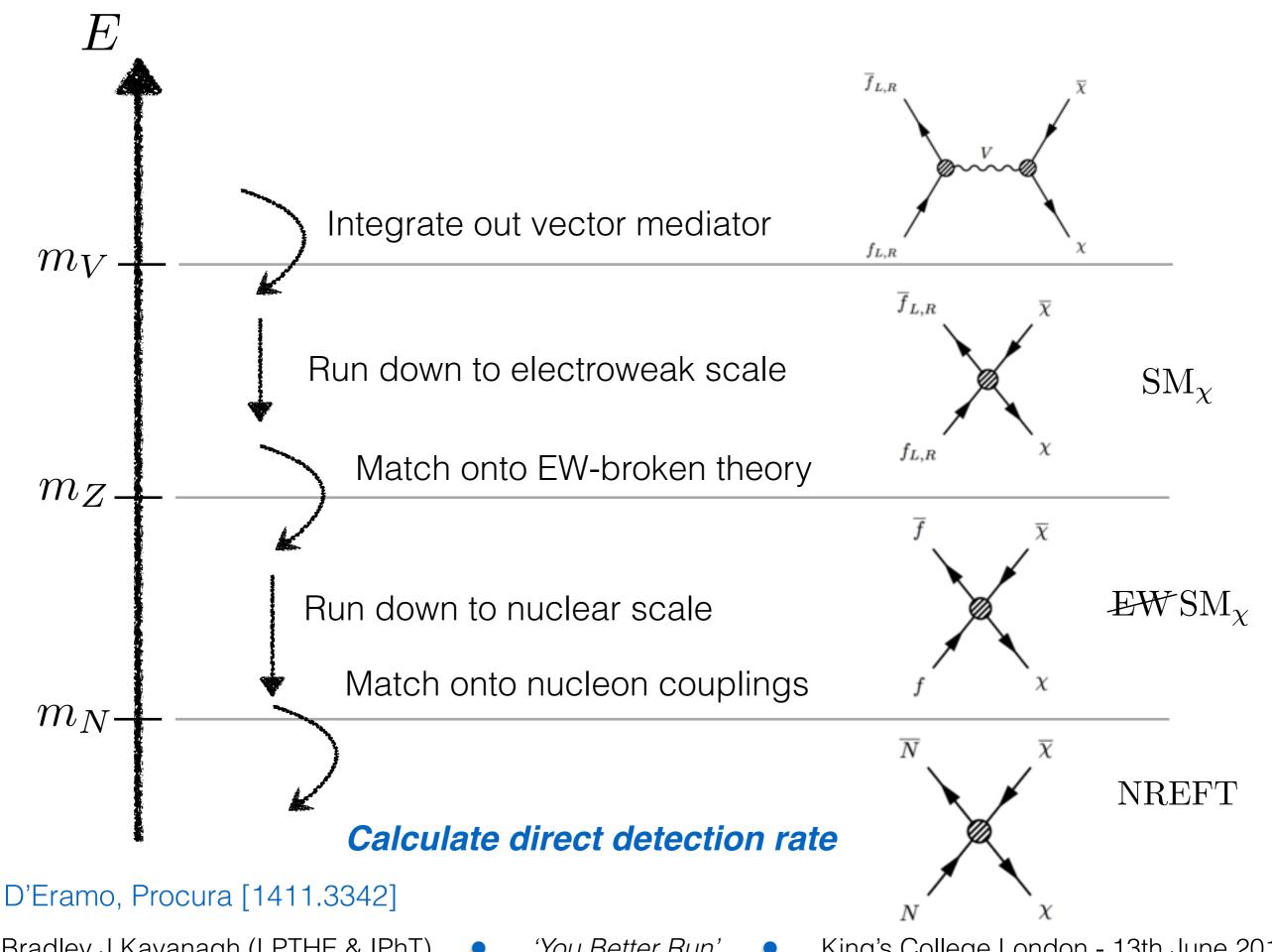
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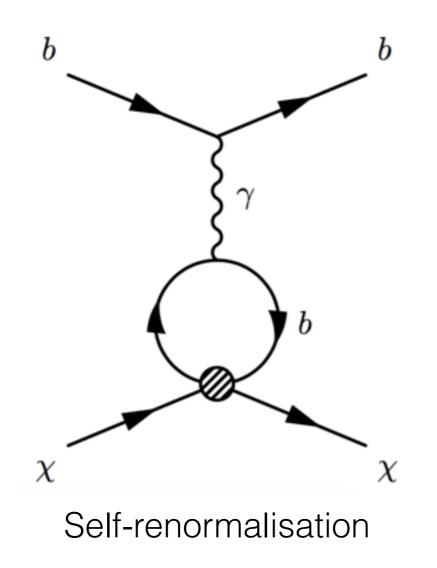


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#### RGE effects

As we move between the different scales, we have to take into account the running of the couplings, due only to loops of Standard Model particles, e.g.

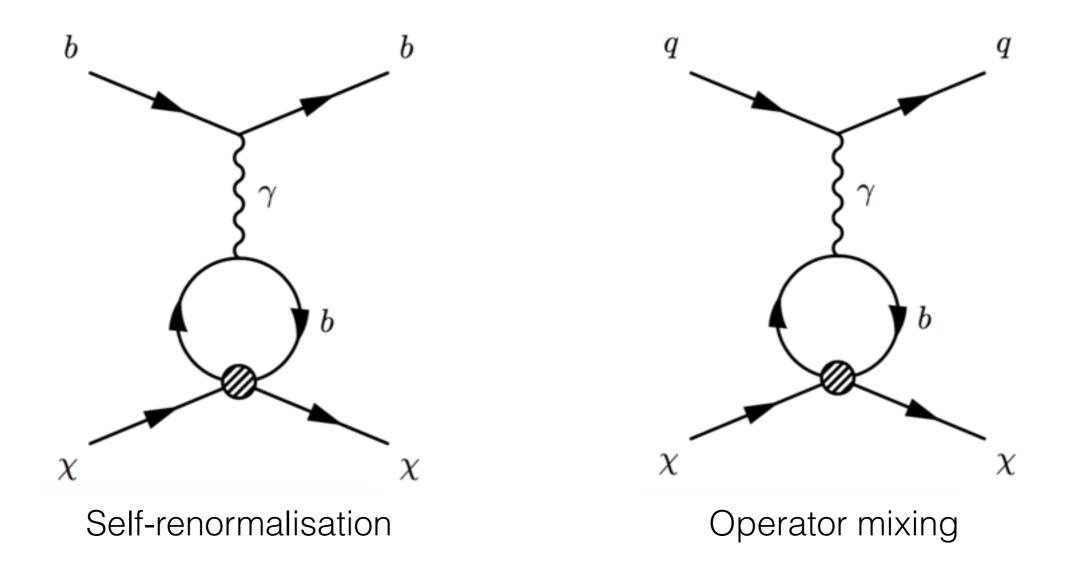


 $\chi$ 

Operator mixing

### **RGE** effects

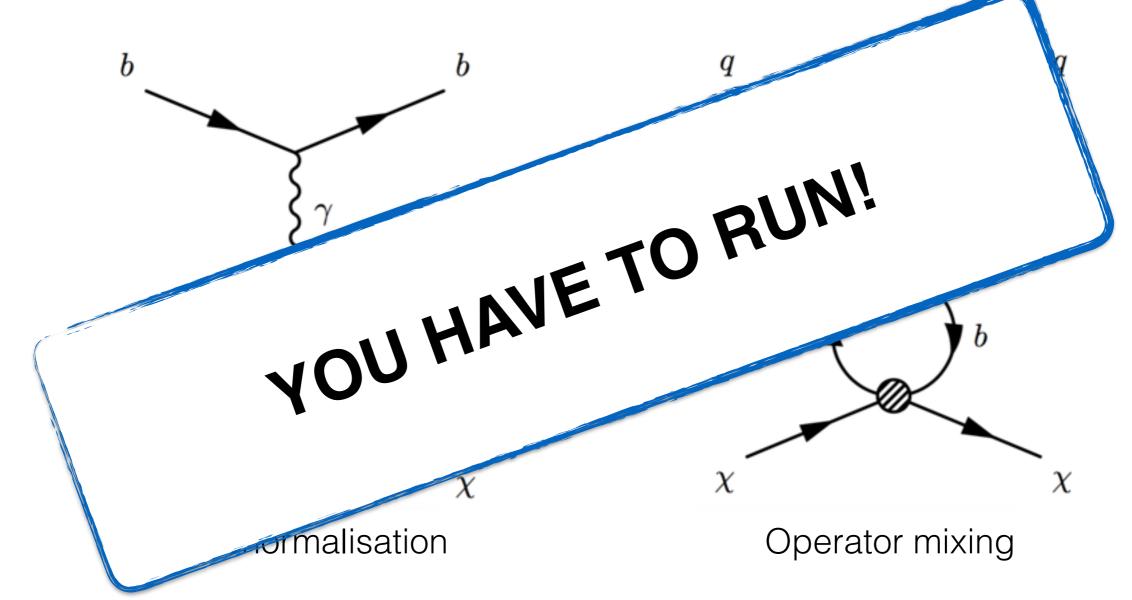
As we move between the different scales, we have to take into account the running of the couplings, due only to loops of Standard Model particles, e.g.



The running doesn't depend on the properties of the Dark Sector.

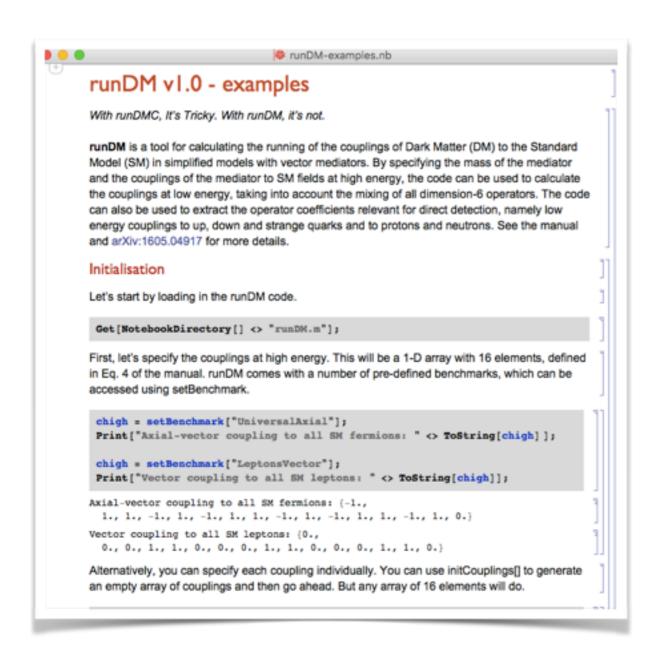
### **RGE** effects

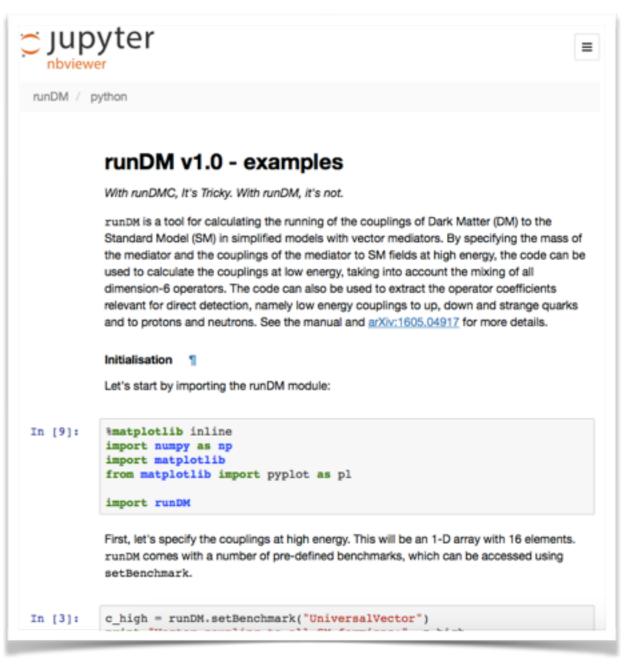
As we move between the different scales, we have to take into account the running of the couplings, due only to loops of Standard Model particles, e.g.



The running doesn't depend on the properties of the Dark Sector.

#### runDM - a code for the RGE





Mathematica and Python versions available at: https://github.com/bradkav/runDM/

### runDM - a code for the RGE

#### Input:

DM-SM couplings at high energy scale

#### **Output:**

DM-SM couplings at another arbitrary energy scale **OR** 

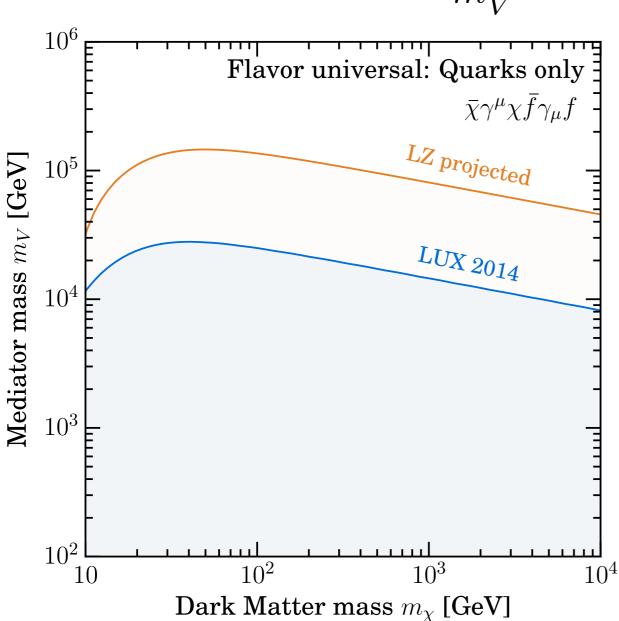
DM-nucleon couplings at direct detection scale

Mathematica and Python versions available at: <a href="https://github.com/bradkav/runDM/">https://github.com/bradkav/runDM/</a>

Direct detection constraints

# Results I - quarks vector

$$\mathcal{L}_{\text{EFT}} = -\frac{1}{m_V^2} J_{\text{DM}\,\mu} \sum_{i=1}^3 \left[ \overline{u^i} \gamma^{\mu} u^i + \overline{d^i} \gamma^{\mu} d^i \right]$$



# Results I - quarks vector

$$\mathcal{L}_{\mathrm{EFT}} = -\frac{1}{m_V^2} J_{\mathrm{DM}\,\mu} \sum_{i=1}^3 \left[ \overline{u^i} \gamma^\mu u^i + \overline{d^i} \gamma^\mu d^i \right]$$
Flavor universal: Quarks only 
$$\overline{\chi} \gamma^\mu \chi f \gamma_\mu f$$

$$LUX \ 2014$$

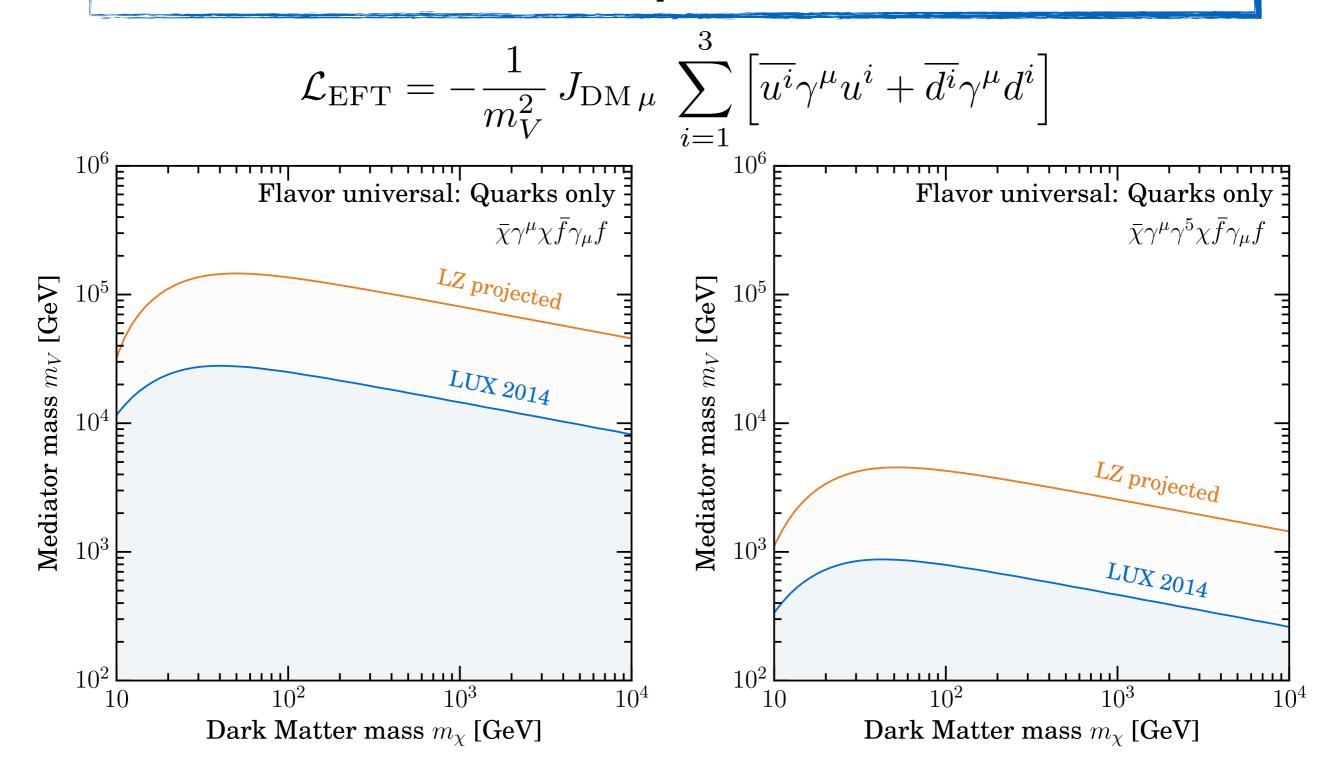
$$Dark \ \mathrm{Matter \ mass} \ m_\chi \ [\mathrm{GeV}]$$

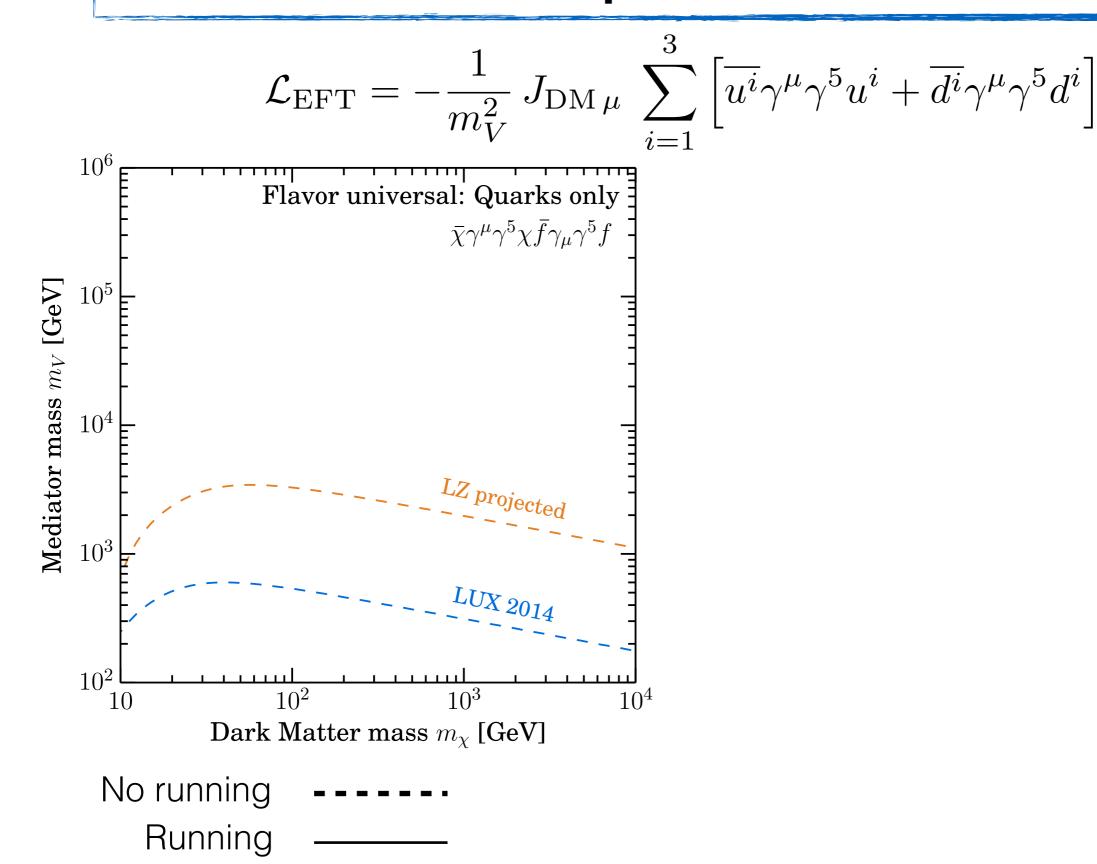
$$c_V^{(q)} \sim 1 + \# \frac{e^2}{16\pi^2} \ln(m_V/m_N)$$





# Results I - quarks vector

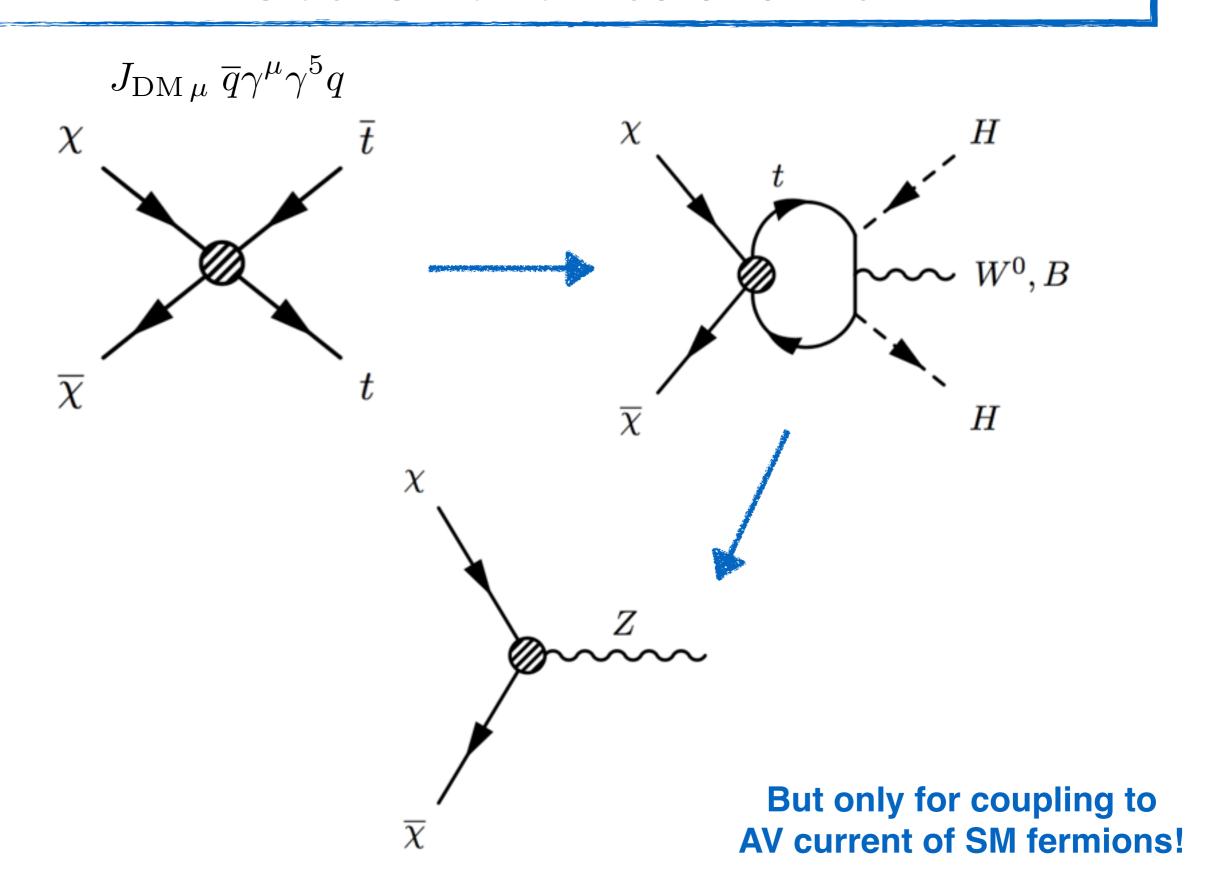




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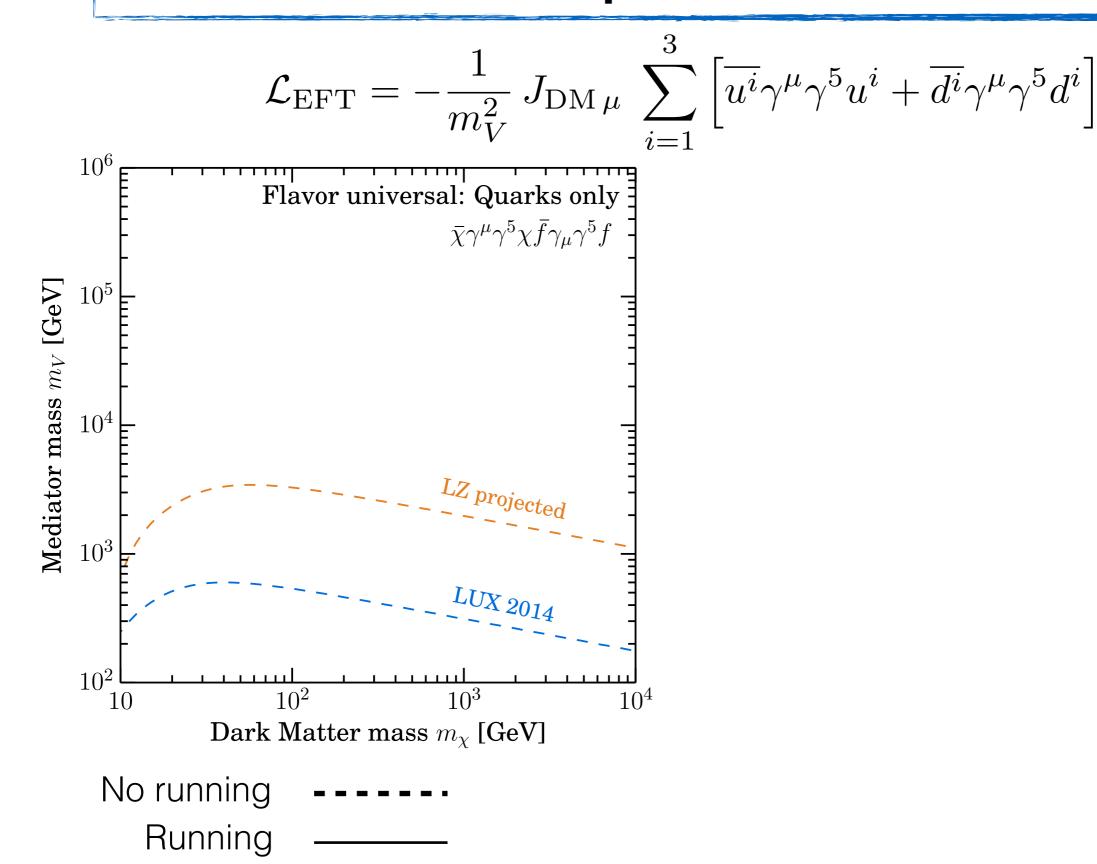
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## Aside: SM axial-vector current



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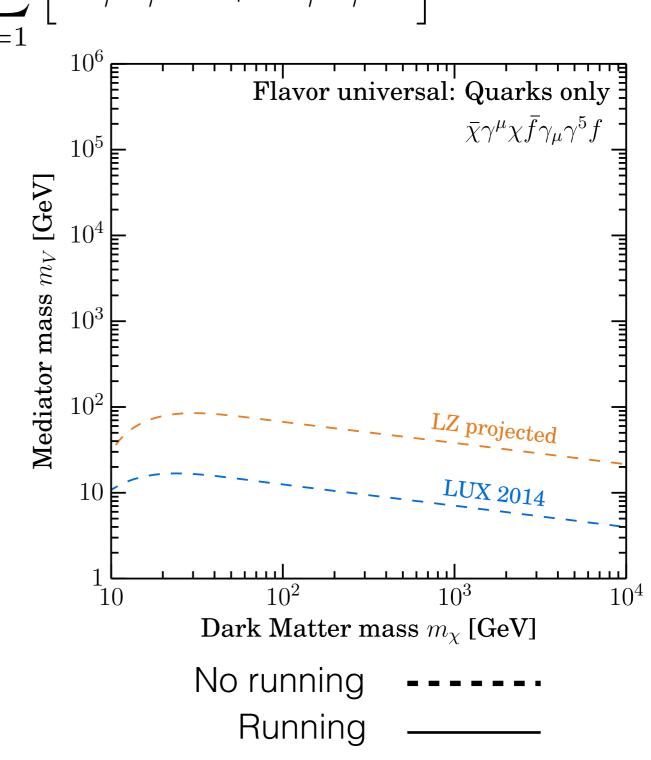
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$$\mathcal{L}_{\mathrm{EFT}} = -\frac{1}{m_V^2} \, J_{\mathrm{DM}\,\mu} \, \sum_{i=1}^3 \left[ \overline{u^i} \gamma^\mu \gamma^5 u^i + \overline{d^i} \gamma^\mu \gamma^5 d^i \right]$$
 Flavor universal: Quarks only 
$$\overline{\chi} \gamma^\mu \gamma^5 \chi \overline{f} \gamma_\mu \gamma^5 f$$
 
$$\overline{\chi} \gamma^\mu \gamma^5 \chi \overline{f} \gamma^\mu \gamma^5 \chi \overline{f} \gamma^\mu \gamma^5 f$$
 
$$\overline{\chi} \gamma^\mu \gamma^5 \chi \overline{f} \gamma$$

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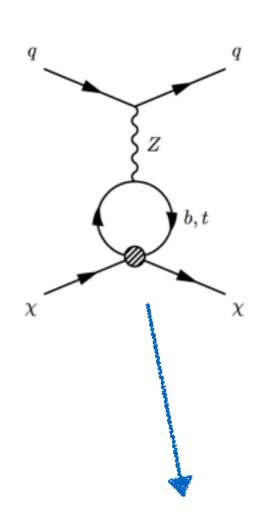
$$\mathcal{L}_{\text{EFT}} = -\frac{1}{m_V^2} J_{\text{DM}\,\mu} \sum_{i=1}^3 \left[ \overline{u^i} \gamma^{\mu} \gamma^5 u^i + \overline{d^i} \gamma^{\mu} \gamma^5 d^i \right]$$



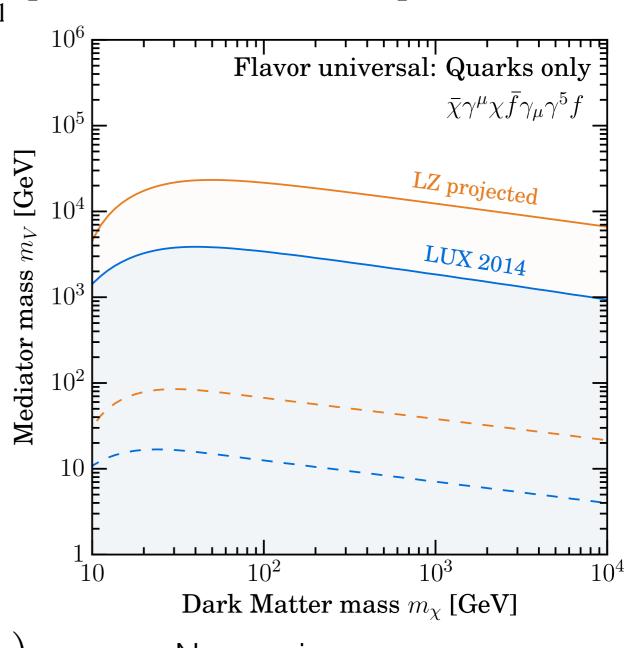
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$$\mathcal{L}_{\text{EFT}} = -\frac{1}{m_V^2} J_{\text{DM}\,\mu} \sum_{i=1}^3 \left[ \overline{u^i} \gamma^{\mu} \gamma^5 u^i + \overline{d^i} \gamma^{\mu} \gamma^5 d^i \right]$$

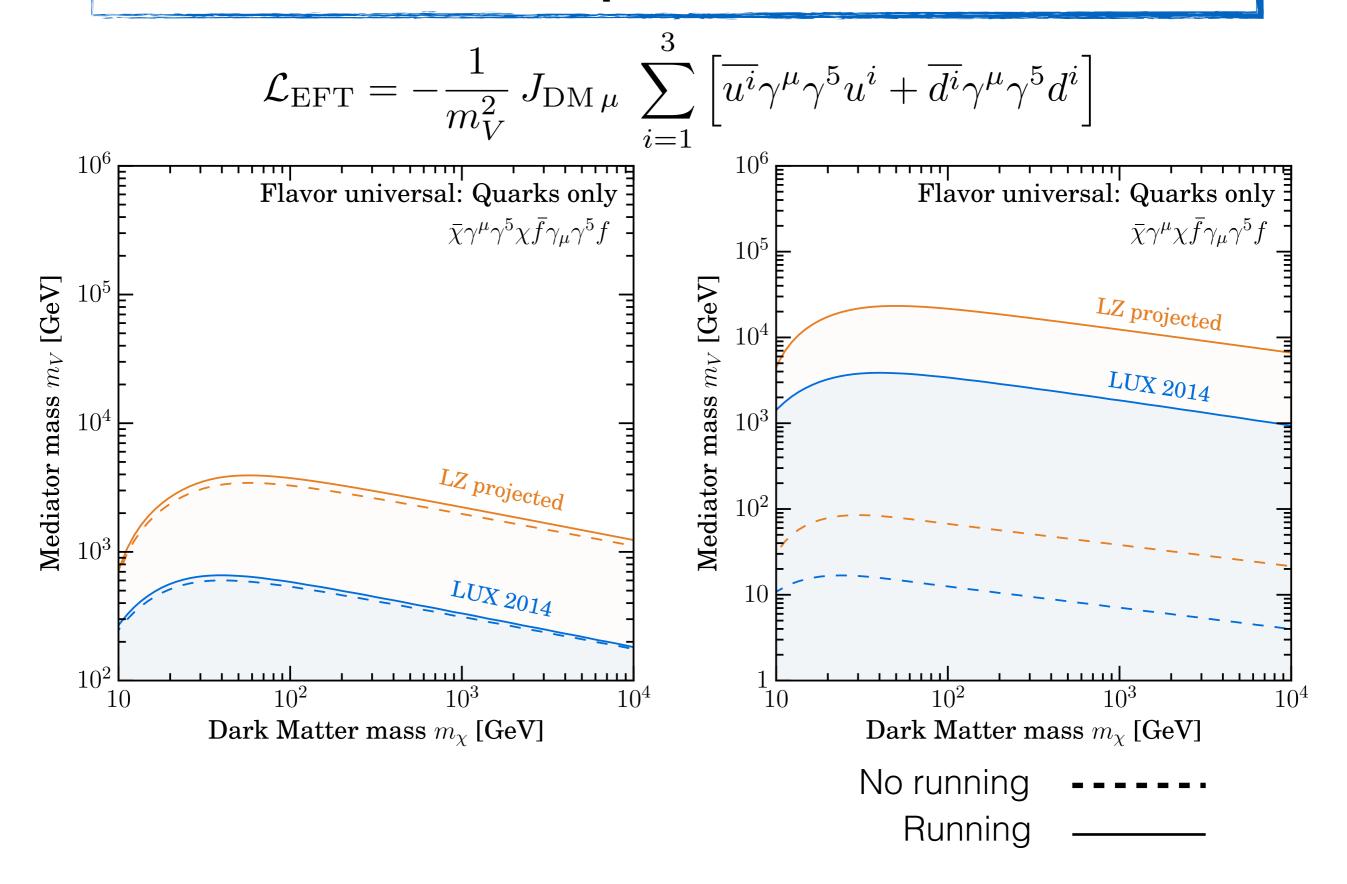


$$c_V^{(q)} \sim 0 + \# \frac{\lambda_{b,t}^2}{16\pi^2} \ln(m_V/m_N)$$



No running ----Running ———

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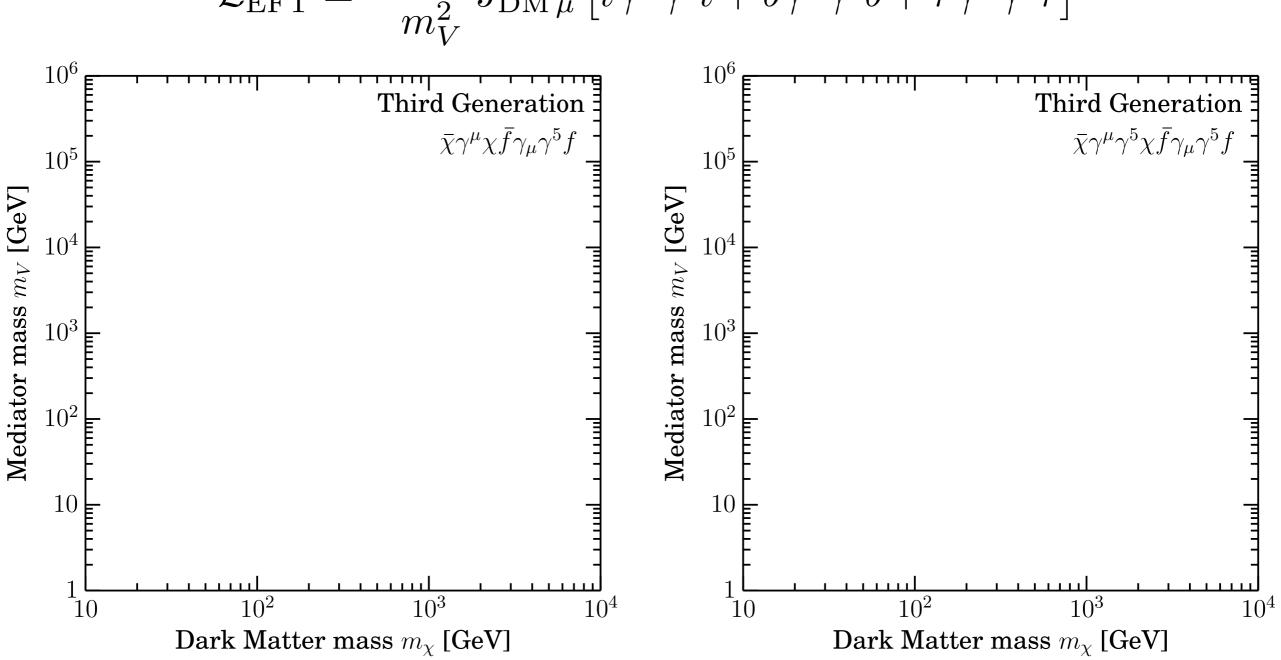


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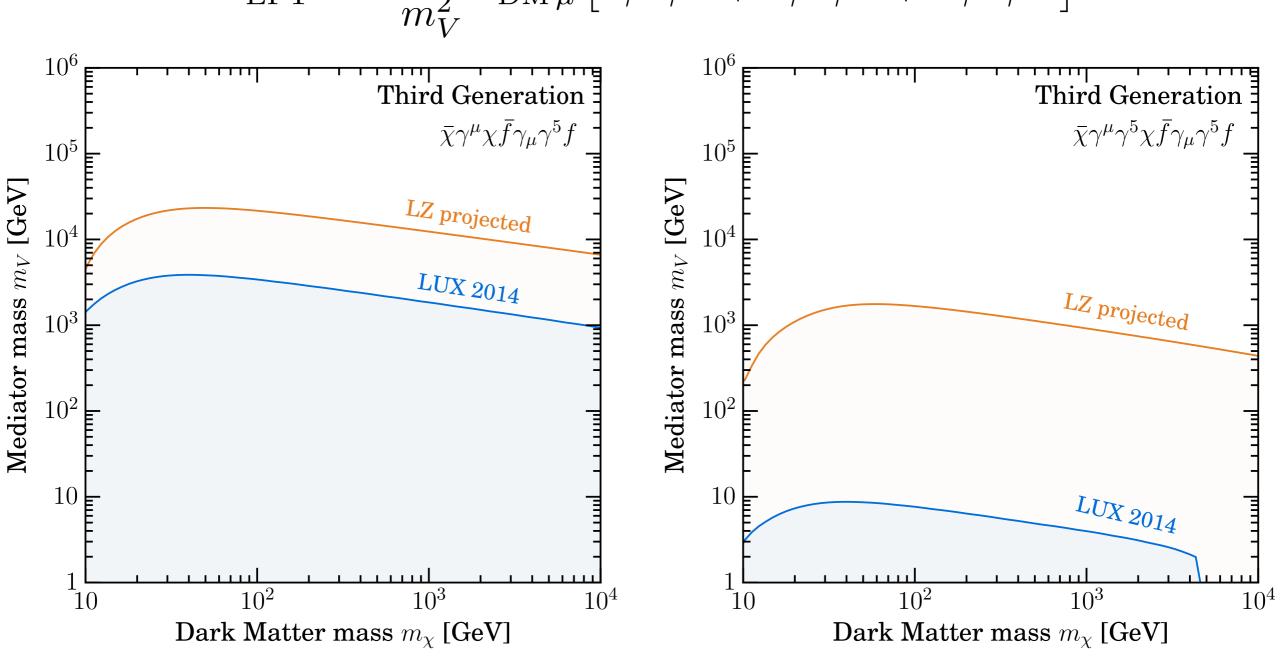
#### Results III - 3rd Generation axial-vector

$$\mathcal{L}_{\text{EFT}} = -\frac{1}{m_V^2} J_{\text{DM}\,\mu} \left[ \overline{t} \gamma^{\mu} \gamma^5 t + \overline{b} \gamma^{\mu} \gamma^5 b + \overline{\tau} \gamma^{\mu} \gamma^5 \tau \right]$$



#### Results III - 3rd Generation axial-vector

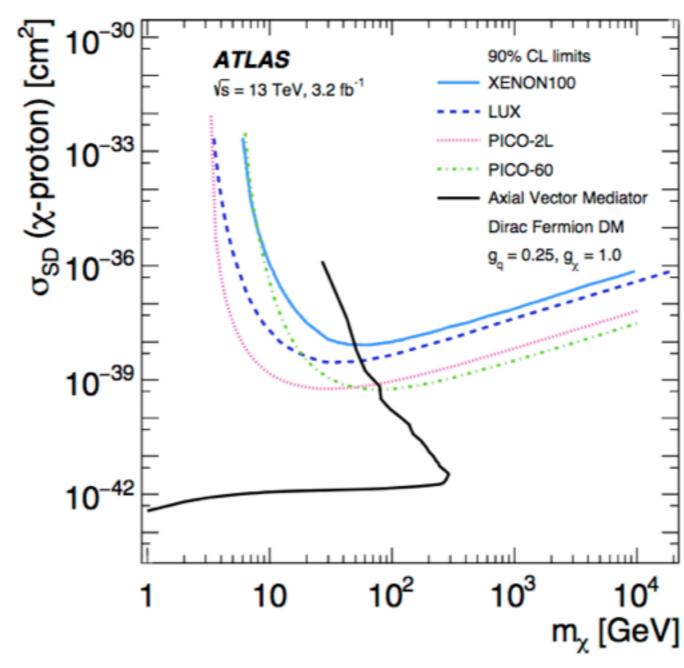
$$\mathcal{L}_{\text{EFT}} = -\frac{1}{m_V^2} J_{\text{DM}\,\mu} \left[ \overline{t} \gamma^{\mu} \gamma^5 t + \overline{b} \gamma^{\mu} \gamma^5 b + \overline{\tau} \gamma^{\mu} \gamma^5 \tau \right]$$

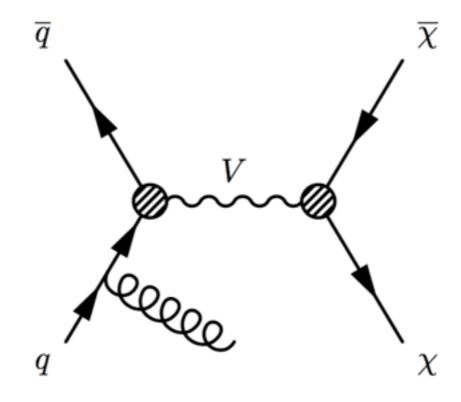


# Comparing DD and LHC searches

#### LHC mono-X searches

$$\mathcal{L}_{AV} = g_{\chi} V_{\mu} \, \overline{\chi} \gamma^{\mu} \gamma^{5} \chi + g_{q} V_{\mu} \sum_{q} \overline{q^{i}} \gamma^{\mu} \gamma^{5} q^{i}$$





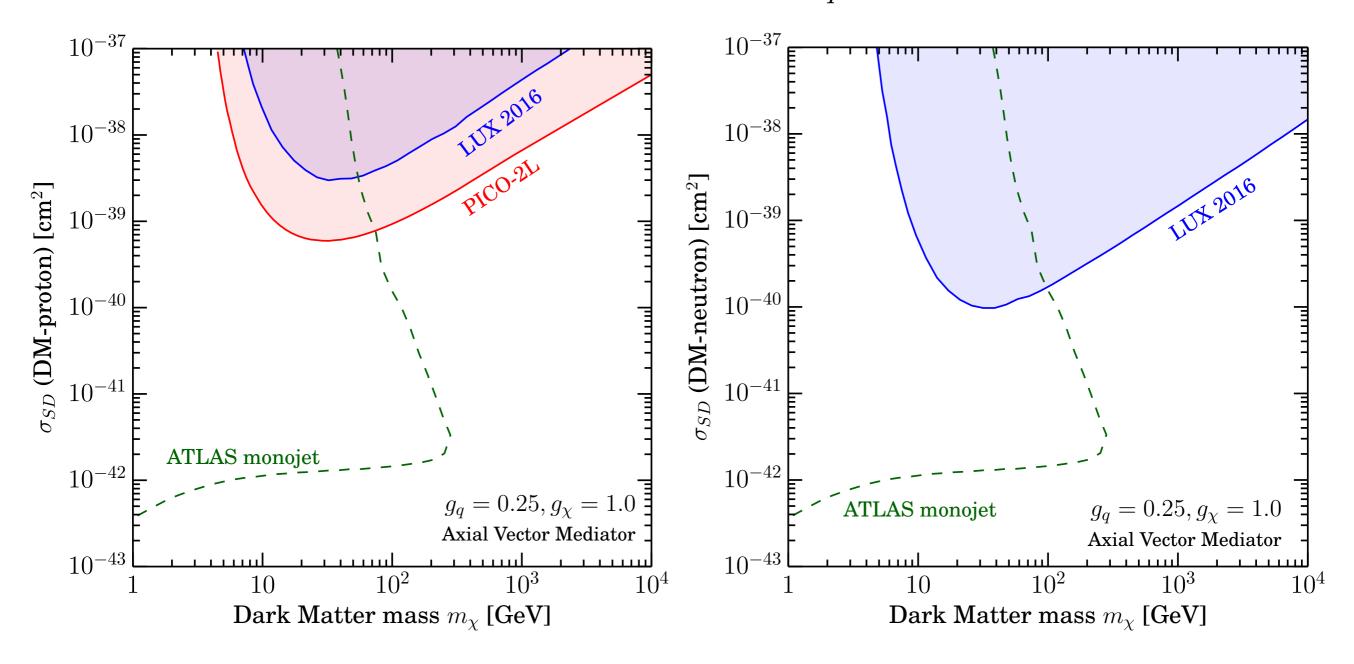
ATLAS [1604.07773]

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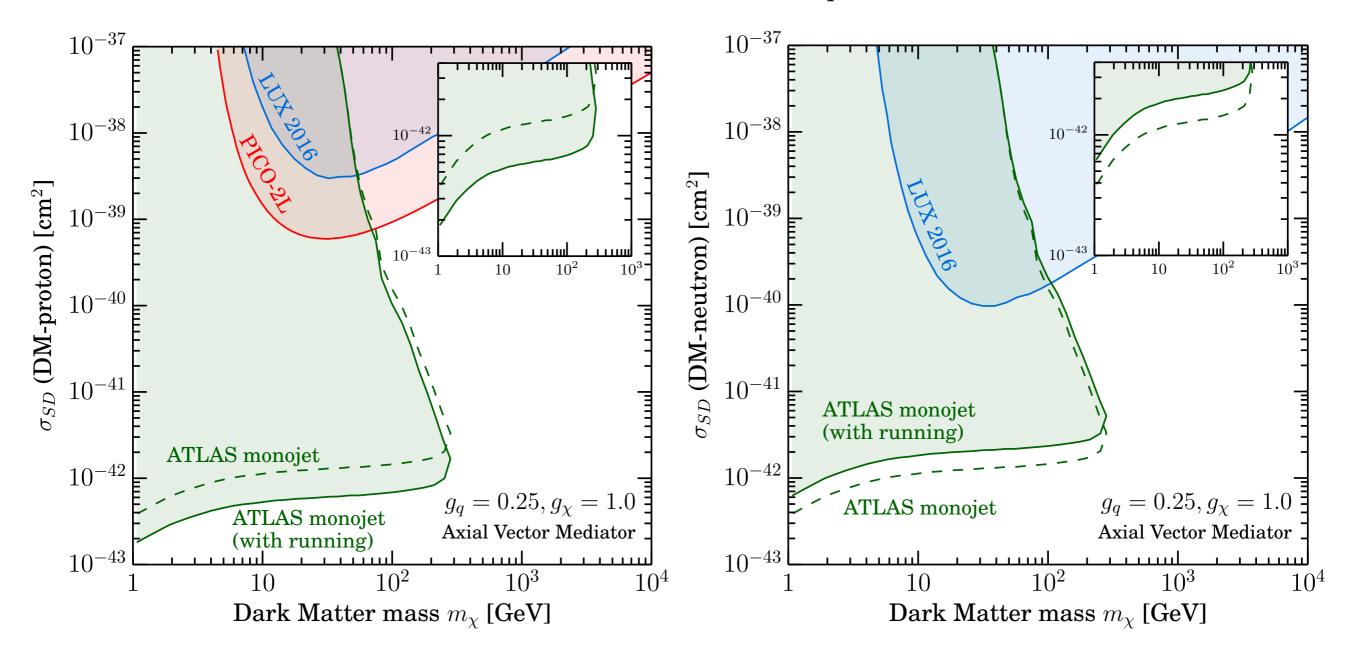
#### LHC mono-X searches

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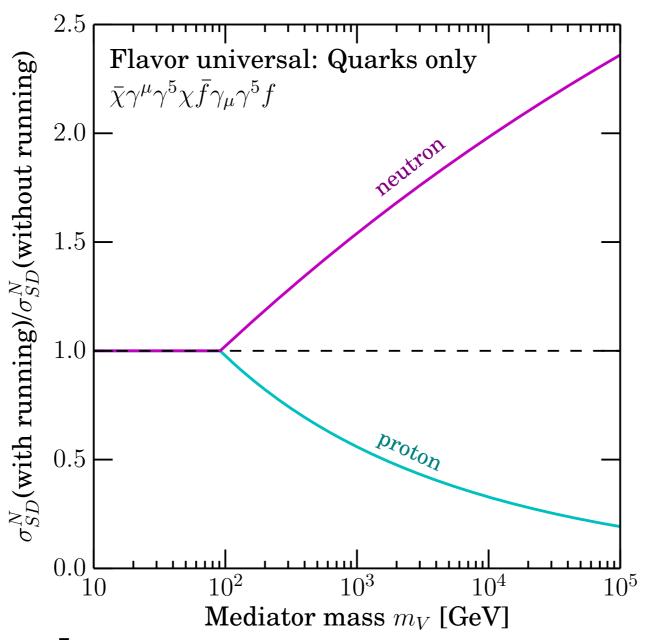
#### LHC mono-X searches

$$\mathcal{L}_{AV} = g_{\chi} V_{\mu} \, \overline{\chi} \gamma^{\mu} \gamma^{5} \chi + g_{q} V_{\mu} \sum_{q} \overline{q^{i}} \gamma^{\mu} \gamma^{5} q^{i}$$



# Isospin violation

$$\mathcal{L}_{AV} = g_{\chi} V_{\mu} \, \overline{\chi} \gamma^{\mu} \gamma^{5} \chi + g_{q} V_{\mu} \sum_{q} \overline{q^{i}} \gamma^{\mu} \gamma^{5} q^{i}$$



$$C_A^{(N)} = g_q \left[ \sum_{q=u,d,s} \Delta_q^{(N)} \right] + \frac{3g_q}{2\pi} \left( \Delta_d^{(N)} + \Delta_s^{(N)} - \Delta_u^{(N)} \right) \left[ \alpha_t \ln(m_V/m_Z) - \alpha_b \ln(m_V/\mu_N) \right]$$

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#### Other interactions

Operator	Coefficient
$ar{\chi}\chiar{q}q$	$m_q/M_st^3$
$ar{\chi}\gamma^5\chiar{q}q$	$im_q/M_st^3$
$ar{\chi}\chiar{q}\gamma^5q$	$im_q/M_st^3$
$\bar{\chi}\gamma^5\chi\bar{q}\gamma^5q$	$m_q/M_*^3$
$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}q$	$1/M_*^2$
$ar{\chi}\gamma^{\mu}\gamma^5\chiar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
$\bar{\chi}\gamma^{\mu}\chi\bar{q}\gamma_{\mu}\gamma^5q$	$1/M_*^2$
$\bar{\chi}\gamma^{\mu}\gamma^5\chi \bar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$
	$ar{\chi}\gamma^5\chiar{q}q$ $ar{\chi}\chiar{q}\gamma^5q$ $ar{\chi}\gamma^5\chiar{q}\gamma^5q$ $ar{\chi}\gamma^5\chiar{q}\gamma^5q$ $ar{\chi}\gamma^\mu\chiar{q}\gamma_\mu q$ $ar{\chi}\gamma^\mu\gamma^5\chiar{q}\gamma_\mu q$ $ar{\chi}\gamma^\mu\chiar{q}\gamma_\mu\gamma^5q$ $ar{\chi}\gamma^\mu\chiar{q}\gamma_\mu\gamma^5q$

Standard SI

Standard SD

Goodman et al. [1008.1783]

## **Other interactions**

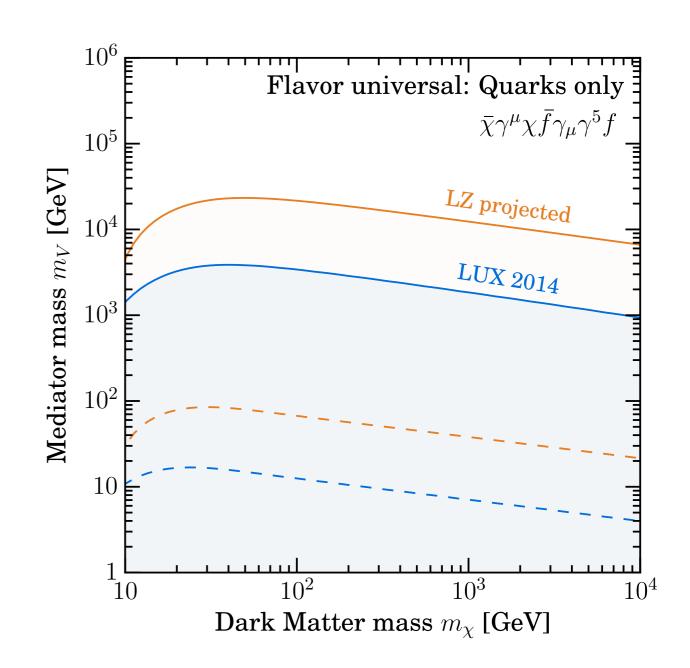
Name	Operator	Coefficient
D1	$ar{\chi}\chiar{q}q$	$m_q/M_st^3$
D2	$ar{\chi}\gamma^5\chiar{q}q$	$im_q/M_st^3$
D3	$ar{\chi}\chiar{q}\gamma^5q$	$im_q/M_st^3$
D4	$ar{\chi}\gamma^5\chiar{q}\gamma^5q$	$m_q/M_*^3$
D5	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D6	$ar{\chi}\gamma^{\mu}\gamma^5\chiar{q}\gamma_{\mu}q$	$1/M_{\star}^2$
D7	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$
D8	$\bar{\chi}\gamma^{\mu}\gamma^5\chi\bar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$

Goodman et al. [1008.1783]

## Other interactions

Name	Operator	Coefficient
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D3	$ar{\chi}\chiar{q}\gamma^5q$	$im_q/M_st^3$
D4	$ar{\chi}\gamma^5\chiar{q}\gamma^5q$	$m_q/M_*^3$
D5	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D6	$\bar{\chi}\gamma^{\mu}\gamma^{5}\chi\bar{q}\gamma_{\mu}q$	$1/M_{*}^{2}$
D7	$ar{\chi}\gamma^{\mu}\chiar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$
D8	$\bar{\chi}\gamma^{\mu}\gamma^5\chi\bar{q}\gamma_{\mu}\gamma^5q$	$1/M_{*}^{2}$

Goodman et al. [1008.1783]



#### **Caveats**

If we add extra degrees of freedom, the phenomenology may be different E.g. Jackson et al. [1303.4717]

May want to include mass and/or kinetic mixing of the mediator

$$\mathcal{L}_{\mathrm{KM}} \sim \epsilon Z_{\mu \nu} V^{\mu \nu}$$
 Langacker [0801.1345]



but we expect this to *strengthen* the limits

In general, we need to worry about the UV completion (e.g. anomaly cancellation, Higgling of the U(1)', etc.)

But if we stick to the Simplified Model framework, our results are valid - and unavoidable!

## Summary

Need to take into account separation of scales



RGE (use runDM code)

Low-E couplings changed and new operators induced



limits on heavy mediators affected (sometimes by OoM!)

Running is important for search complementarity



req. for translating LHC searches into DD plane

Isospin violation (factor of ~few) is automatic



arising only from SM loops

# **Summary**

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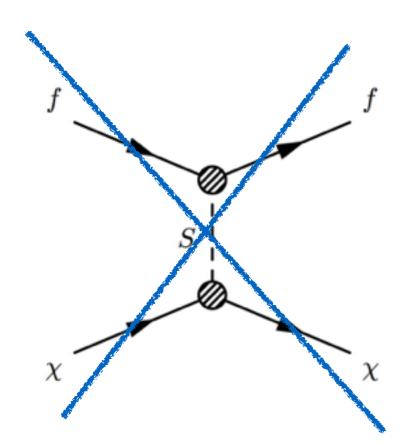
YOU HAVE TO RUN!

# Backup slides

#### **Scalar Mediators**

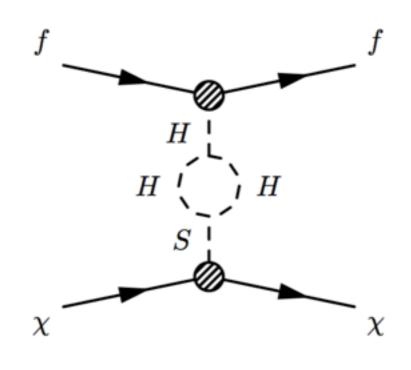
Interactions through a scalar mediator appear at dimension-7, with rates typically suppressed by the quark mass

$$O^S_{gg} = rac{lpha_s}{\Lambda^3}\,ar\chi\chi\,G_{\mu
u}G^{\mu
u}\,,\quad O^{SS}_{qq} = rac{m_q}{\Lambda^3}\,ar\chi\chi\,ar q q\,,$$



Crivellin, D'Eramo, Procura [1402.1173]

Buckley et al. [1410.6497]



#### Scalar mediator: 750 GeV

# A 750 GeV Portal: LHC Phenomenology and Dark Matter Candidates

Francesco D'Eramo a,b, Jordy de Vries c, Paolo Panci d

D'Eramo, de Vries, Panci [1601.01571]

$$\mathcal{L}_{\mathrm{EFT}}^{m_S < \mu < \Lambda} = \sum_{q=u,d,s,c,b,t} \frac{c_{yq} y_q}{\Lambda} S \left( \overline{q}_L H q_R + \mathrm{h.c.} \right) + \frac{c'_{GG} \alpha_s}{\Lambda} S G^{A \mu \nu} G^A_{\mu \nu} ,$$

$$\mathcal{C}_q(\mu_N) \simeq -5.86 \, \mathcal{C}_{GG}(m_S) ,$$

$$\mathcal{C}_{GG}(\mu_N) \simeq 4.01 \, \mathcal{C}_{GG}(m_S) .$$

Substantial RG effects!

'You Better Run'

# The operator basis

$$rac{J_{{
m DM}\,\mu}\;H^{\dagger}\,i\overleftrightarrow{D}^{\mu}H}{m_{V}^{2}}$$

$$\frac{J_{{\rm DM}\,\mu}\,\partial_{\nu}B^{\mu\nu}}{m_{V}^{2}}$$